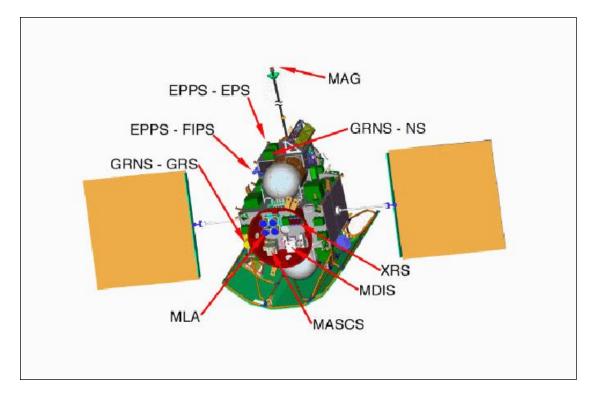
MESSENGER:

SoftwareInterfaceSpecificationforthe DerivedDataRecordsoftheEnergeticParticle andPlasmaSpectrometer

Version1M



Preparedby

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Document Review

This document and the archive it describes have been through PDS Peer Review and have been accepted into the PDS archive.

George Ho, MESSENGERE PPS Instrument Scientist, has reviewed and approved this document.

Jim Raines, MESSENGER FIPS Instrument Scientist, has reviewed and approved this document.

Susan Ensor, MESSENGER Science Operations Center Lead, has reviewed and approved this document.

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Table 1 Revision History

Version	Author	Date	Description	Sections
1A	L.Nguyen	5/22/2012	Initialrevision	All
1B	M.Gannon	5/25/2012	MinorupdatesfromAPLteamreview	All
1C	J.Raines M.Gannon	10/17/2012	UpdatesfromtheEPPSPDSPeerReviewofthe SampleArchiveforDDR	All
1D	P. Bedini	1/29/2013	Editorialupdates	All
1E	M.Reid	1/29/2013	Replaced text in dicating a draft version of this document with approved & release version.	Document Review,pg. 2
1F	J.Raines M.Gannon	11/06/2013	UpdatestoincorporatenewFIPSAdvanced Products	
1G	G.Ho J.Raines S.Ensor	1/17/2014	Updatesto DDRdescriptions andapplicable documents	1.2, 2, 5.2.1
1Н	J.Raines M.Reid	6/13/2014	AddedinformationabouttheFIPSViewing Normalizationproducts; fixedlargerdocument formatting.	5.2.2, 5.3.4,6.5, 6.8, 8.7, 8.8
11	M.Reid	7/16/2014	Removedreferencestothe FIPS PixelFOVtable. TheyweremovedtotheCDRSIS. Corrected documentvolumedirectorystructurediagram.	5.2.2.6, 5.3,6.5, 6.7,8
1J	J.M.Raines M.Reid	11/19/2014	Modifiedtoincorporate substitution of FIPS_ARRDIR productwith FLUXMAPand additionofERPCHANGproduct. Added ERPCHANGproduct.Rem ovedFIPS_FOVPIX directory (FIPA_*filesmovedtoDocumentVol. CALIBRATION). ModificationstoROTSMOlabel. Formatting.	5.1.1, 5.2,5.3. 4, 6.8,8.3, 8.7.

Version	Author	Date	Description	Sections
1K	J.M.Raines M.Gannon	1/27/2015	EditstoFIPSFluxMapandFIPSPitchAngle productdescriptions. Correctionofminor typographicalerrors.	5.2.2.3, 5.2.2.5
1L	S.Ensor	1/16/2016	Finaledits(partial)	All
1M	J.M. Raines M.Gannon	1/28/2016	Finaledits	1.1.4and others

1 Purpose and Scope of Document

1.1 Purpose

This documents ervest to provide users of the MErcury Surface, Space EN vironment, GE ochemistry, and Ranging (MESSENGER) Energetic Particle and Plasma Spectrometer (EPPS) data products with a detailed description of the EPPS instrument, data product generation, validation, and storage Note that the EPPS is made up of two instruments ubsystems, the Fast Imaging Plasma Spectrometer (FIPS), and the Energetic Particle Spectrometer (EPS) . The FIPS and EPS are described in individual sections within this document. They are referred to separately when necessary and referred to as the EPPS instrument when dealing with a reas common to both instruments. The FIPS covers the lower energy range of particles and measures the mass percharge (m/q), energy percharge (E/q), and incoming direction of each charged particle. The EPS covers the higher energy range and measures mass, energy, and incoming direction of each particle. The MESSENGEREPPS data products a redeliverable sto the Planetary Data System (PDS) and the scientific community that its upports. All data for mats are based on the PDS standard.

1.2 Scope

The EPPS science data are divided into two categories: Level 2 edited raw data (referred to as experiment data records or EDRs) and processed data (referred to as reduced data records or RDRs). RDRs are generated from EDRs, and represent data calibrated to aphysical unit such as particle intensity (Level 3), resampled Level 4 data products, or derived Level 5 data products. RDRs consist of Calibrated Data Records (CDRs), Derived Data Records (DD Rs), and Derived Analysis Products (DAPs). This Software Interface Specification (SIS) describes the EPPSDDR data products.

For EPS, DDRs consist of two products: pitch angle values, and pitch angle distribution spectrograms over selected ranges of energies for selected time periods. For FIPS, DDRs consist of seven products that provided at a for spatial and temporal distributions of observed density for majorion species, and for selected ion species and time periods, energy spectra, pitch angle distributions, ion flux, density, temperature, and pressure, and viewing normalizations for each energy scan . The DDR data were delivered to the PDS as CODMAC (Committee on Data Management and Computation) Level 4 data. EPPS's DDRs are formatted to include standard PDS labels. A detailed description of all data products in the EPPS DDR follows.

In addition, this SIS describes the EPPS document at ion volume, which contains products related to both the EDR- and R DR-level archives. The contents of the document at ion volume enables one to the contents of the document at ion volume enables one to the contents of the document at ion volume enables one to the contents of the document at ion volume enables one to the contents of the conte

conductusefulanalysisoftheDDRs.Thedocumentationvolumeisdescribedingreaterdetailin section 6.6.

2 Applicable Documents

 $The {\tt MESSENGEREPPSSIS} is responsive to the following documents:$

- 1. PlanetaryDataSystemStandardsReference,Feb27,2009,Version3.8.JPLD-7669,Part-
- 2.
- **2.** MESSENGERDataManagementandArchivingPlan.TheJohnsHopkinsUniversity,APL. DocumentIDnumber7384-9019
- **3.** MESSENGERMercury:Surface,SpaceEnvironment,Geochemistry,Ranging;Amissionto OrbitandExplorethePlanetMercury,ConceptStudy,March1999.DocumentIDnumber FG632/99- 0479
- **4.** [PLR]Appendix7tothediscoveryprogramPlan;ProgramLevelRequirementforthe MESSENGERDiscoveryproject;June20,2001.

The following documents may be referenced for details on the EPPS instruments:

- **5.** Livietal.(Theenergeticparticlespectrometer(EPS)onMESSENGER:Instrument description,characterization,andcalibration,MESSENGERProjectreport,2004)
- **6.** Zurbuchenetal.(TheFastIonPlasmaSpectrometer(FIPS)calibrationreport,MESSENGER Projectreport,2004)
- **7.** Andrewsetal.(TheEnergeticParticleandPlasmaSpectrometerInstrumentonthe MESSENGERSpacecraft,SpaceScienceReviewsVolume131,Numbers1-4,August2007)
- **8.** Raines et al. (Distribution and compositional variations of plasma ions in Mercury's space environment:ThefirstthreeMercuryyearsofMESSENGERobservations,Journalof GeophysicalResearch,118,p1604-1619,2013)
- **9.** Gilbertetal.(Backgroundnoiseinspace-basedtime- of-flightsensors,ReviewsofScientific Instrumentation,inreview,2014)

- **10.** Gershmanetal.(Post-processingmodelingandremovalofbackgroundnoiseinspace-based time-of-flightsensors,DeepBlue,2013, http://hdl.handle.net/2027.42/100358)
- **11.** Rainesetal.(MESSENGERobservationsoftheplasmaenvironmentnearMercury,Planetary andSpaceScience59,p2004-2015,2011)
- **12.** Ho et al. (Spatial distribution and spectral characteristics of energetic electrons in Mercury's magnetosphere, *J.Geophys.Res.*,doi:10.1029/2012JA017983,117,2012)
- **13.** Gershmanetal.(MagneticFluxPile- up and Plasma Depletion in Mercury's Subsolar Magnetosheath, JournalofGeophysicalResearch:SpacePhysics, 118, p7181-7199, 2013)
- **14.** Slavin et al. (MESSENGER Observations of Extreme Loading and Unloading of Mercury's MagneticTail,Science,329,p665-668,2010,doi:10.1126/science.1188067)

3 Relationships with Other Interfaces

The EPPSDDR data products were stored at the MESSENGERS cience Operations Center (SOC) during the MESSENGER mission. The data products were transferred to the PDSP lanetary Plasma Interactions (PPI) Node according to the delivery schedule in section 7. The data in the DDR files is stored in PDSASCIITABLE objects unless stated otherwise (section 5.2).

4 Roles and Responsibilities

Therolesandresponsibilitiesoftheinstrumentteams, The Johns Hopkins University Applied Physics Laboratory (APL), Applied Coherent Technology Corporation (ACT), and the Planetary Data System (PDS), are discussed in sections 5.3.2 and 5.3.3.

5 Data Product Characteristics and Environment

5.1 Instrument Overview

5.1.1 FIPS Overview

The Fast Imaging Plasma Spectrometer (FIPS) measures the energy per charge (E/q), time- of-flight (TOF), and incident angles for plasma ions entering the sensor . Intensities, velocity distributions, and mass per charge (m/q) distributions are derived from the seme as ure ments and make up FIPS primary science data . These data are used to understand the kinetic properties, angular distributions, and composition of Mercury magnetos phericions, and contribute to the characterization of the planetary magnetic field.

Ions measured by FIPS pass through an electrostatic analyzer (ESA), located at the entrance to the sensor; apost-acceleration chamber between the output of the ESA and the carbon foil; and at ime-of-flight telescope. The ESA at the entrance to FIPS acts as a wide-angle lens for ions, with a 1.4 sr field of view. It allows only ions with a specific E/q band to enter through its output plane and is stepped through 64 values to complete one measurement cycle (scan), no minally from 0.046-13.3 keV/e. FIPS is normally operated in one of two stepping rates: one stepper second (normal mode), or one stepper 100 milliseconds (burst mode). When delays due to high-voltage ramp-ups are included, these resultincy cletimes of 64 secand 8 sec, respectively arbitrary values, different for each step and in the should, as the should, as ettling time, and an integration (dwell) time and an integration (dwell) time.

Ions exit the output plane of the ESA and are then accelerated in the post acceleration chamber. This acceleration is done to give low-energy ions sufficient energy to penetrate the carbon foil . The acceleration also helps to reduce energy straggling and angular scattering -effects that cause degradation in mass resolution and imaging . When ions exit the carbon foil, secondary electrons are liberated. These electrons travel to the Start MCP (microchannel plate), providing a timing-start signal and incident angle information via impact location on a position-sensing anode. The ion then travels through the TOF chamber and strikes the Stop MCP, providing a timing-stop signal and allowing computation of TOF . From E/q and TOF, m/q can be computed . FIPS can measure species from HtoFe, 1-60 amu/e (or higher).

IndividualionsareidentifiedinFIPSdatafromtheir E/qandTOFmeasurements . Asimplemodelis usedtopredicttheTOFrangeexpectedforeachionasafunction of E/q, referredtoasthe E/q-TOF trackforthation. This model includes the effects of energy lost upon passage through the carbon foil, as well delay squetoelectron flight time and electronic processing. All events with measured TOF within the predicted range for a particular ion are assigned to that ion. To improve signal to noise, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some ions are grouped together. In this dataset, E/q-TOF tracks for some

FIPS uses a double coincidence technique to greatly reduce background noise. However, spurious double-coincidence counts still dooc cur. These counts come from two mains our ces: the extension of very high count proton measurements into other times of flight, and the release of small numbers of the count of the co

ionsfromsurfaceprocesseswithintheinstrument[9].Whileallmajorionspeciesreportedhere canbeanalyzedfromtherawdata,accuracyissignificantlyimprovedbyremovingthesespurious counts.Adetailednoisemodelandremovalmethodhasbeendeveloped[10]andisemployedon thedatainth isworkattheindividualscanlevel .

Afterspeciesareidentifiedin E/q-TOFspaceandnoisecountsremoved,theresultingcountsfor speciess ($C_{i,s}$)ateachmeasured(E/q)_i are transformed to phase spaced ensity ($f_{i,s}$) in units of $f_{i,s}$ using the CDR fluxes ($f_{i,s}$) in sec $f_{i,s}$ using the CDR fluxes ($f_{i,s}$) in sec $f_{i,s}$ using the following relationship:

$$f_{i,s} = J_{i,s} \frac{m_s}{v_{i,s}^2 C}$$

where $v_{i,s}$ is velocity, m_s is ion mass (in a mu) and C is a unit conversion factor, 1.6022 × 10 $^{-20}$ (cm 2 s keV) $^{-1}$.

The details of the conversion of counts to phase space density are provided in the EDR2CDR document, available in the CALIBRATION directory of the EPPS document at ion volume.

DetailsofFIPSoperationscanbefoundin[7].

5.1.2 EPS Overview

 $EPS is a compact TOF teles cope with two main components: a TOF section and a Solid State Detectors (SSD) array. The SSD array comprises sixion-implanted planar silicon detectors, each with four pixels (two dedicated to ion measurements and two to electron measurements) for a total of 24SSD elements. Particles enter the system through a mechanical collimator that delimits the look direction into the instrument and particles that pass through the collimator then transit through a thin, composite Startfoil (polyimide+aluminum, 10 <math display="inline">\mu g/cm^2$) and onto the TOF region of the instrument.

ElectronsarereleasedfromtheinnersurfaceoftheStartfoilandfocusedtoawell-definedregion on an MCPtogeneratetheSTARTsignalinadedicatedanode.Theincidentionsarenot significantlyaffectedbytheelectricfieldsofthefocusingoptics.After a 6cmflightpath,ions

traversetheStopfoil,whichisapolyimide+palladium(19 $\mu g/cm^2$)compositefoil.Thesecondary electronsreleasedbythestopfoilaresteeredtotheMCPandgeneratetheSTOPsignal.Electron trajectorysimulationsshowthatthereislessthan2nsdispersioninthetransittimeofthe secondaryelectronfromthefoiltotheMCP.Sub-nanoseconddispersionisrequiredsoasnotto misidentifyionspecies.IfbothaSTARTandSTOPsignal(doublecoincidence)areregistered,then thetime,t, fortheparticletotravelaknowndistance(d=6cm)canbedetermined.Fortriple coincidencewemusthavetheSTART,STOP,andalsotheenergymeasured(E_{meas})bytheSSDs . Usingthesemeasuredparameters,wecancalculatethemass(M)andtheincidentenergy(E)of eachionusingthefollowingequations:

$$E_{true} = E_{meas} / \gamma$$

$$E_{ToF} = \beta E_{true}$$

$$M = 2(T/d)^2 E_{ToF} = 2(t/d)^2 (\beta E_{meas} / \gamma)$$

 E_{true} takes into account the smallener gyloss of the ions passing through the stop foil, and the energy loss and pulse-height defect in the SSDs. E_{ToF} takes into account the even smaller energy loss or gain in the start foil, and may also include up to ~2.5 keV electrostatic pre-acceleration of ions that remain charged on exiting the start foil . If the energy of the incident particle is not large enough to trigger the SSD, the nonly t is measured and the pulse height of the start anode is used to discriminate whether it is a light (M ~1 amu), or heavy (M >1 amu) ion. At the same value of TOF, heavy ion shave been shown to generate substantially more secondary electrons than do protons.

Besides composition measurements, the particle's angular direction can be determined. The pair of startands to panodes provide the polar entrance angle of the incident particle. The polar angle of $+80^{\circ}$ to -80° is divided into six equals ectors (no minally 27°).

5.1.2.1 Electronmeasurements

 $Energetice lectrons have higher penetration power than ion satthesame energy and in the SSDs dedicated to electron detection in EPS are covered by a thin layer (flashing) of 1 $$\mu$ mo faluminum. This dead layers to psprotons with energy less then $$\sim 250 \text{keV}$; Electrons, on the other hand, lose less than 10 keV by the interaction with this dead layer. Electrons are identified in EPS by the presence of an energy signal. The TOFs pectra collected in the adjacent SSD (without flashing) are used during ground data analysis for checking and correcting for the proton contamination.$

"Calibration" for a particle instrument like EPS means determining the following:

- 1. Transferfunctionfromcountsintoflux(physicalunits)
- 2. Characteristic of "Rate-out" versus "Rate-in"
- 3. Responsetolowenergyandhighenergyparticlebackground
- 4. Responsetovisibleandultravioletlight
- 5. Responsetohighmagneticfield

All these functions need to be characterized and the relevant parameters need to be determined before flight.

5.1.2.2 TransferFunction

Flux, differential intensity, and phases pacedensity

 $The number of particles Nthattraverse an area Aduring a time tcan be characterized by the flux \cite{thm2/s}$

F

N=A*t*F

orbytheintensityI[1/cm2/s/sr]

N=A*t* $\int I\cos(\hat{u})d\Omega$

where Ω is solid angle and Ω is an algorithm and in the sensor determine the limits on the integration.

Oftenusedisthequantitydifferentialintensityf[1/cm2/s/sr/keV],definedasthenumberof particles with energy between E and E+ Δ E that traverse the area A during the timet,where

$$N(E)=f(E)*A*t*$$
 $\Delta\Omega * \Delta E$

In three dimensions, with θ being the polar angle and ϕ the azimuthal angle of a polar reference system:

$$d3N(E,\theta,\phi)=f(E,\theta,\phi)*A\cos(\hat{u})*t*dE\cos\theta d\theta d\phi$$

Note that $f(E,\theta,\phi)$ is related to the phase-spacedensitypsd(number of particles in the configuration space element d3R and with velocity between vand v+d3v) by the simple relationship in the non-relativistic limit (valid for ions measured by EPS but not for the higher energy electrons):

$$psd(s3/cm6)=f(1/cm2/s/sr/keV)*m/v2$$

For relativistic particles, one generally utilizes momentum spacerather than velocity space, and the corresponding expression is:

$$psd(s3/gm.cm6)=f(1/cm2/s/sr/keV)/p2$$

Where "p" is momentum.

Definition of sensor transfer function and geometric factor

Thenumber of counts No fparticles of mass m, in the energy band around mean energy E, angular band $\Delta\theta$ around mean polar direction θ , and angular band $\Delta\phi$ around the mean azimuthal direction ϕ , measured by the instrument during the time δt can be expressed as:

$$N(E, \theta, \phi; m) = \delta t * \int \Delta E \int \Delta \theta \int \Delta \phi f(E, \theta, \phi; m) * A \cos(\hat{u}) * dE \cos \theta d\theta d\phi$$

If $f(E, \theta, \phi)$ is a Dirac δ function (monoenergetic, infinitely narrow beam), then

$$N(E, \theta, \phi; m) = \delta t * f(E, \theta, \phi; m) * G(E, \theta, \phi; m)$$

Where $G(E, \theta, \phi; m)$ [cm2 sr keV] is the transfer function of the instrument.

Intheotherlimit, when the flux is completely isotropic (all directions the same)

$$N(E; m) = \delta t * f(E; m) * GF(E; m)$$

GF is called geometric factor and represents a measure of the efficiency of the system (count rate/flux), and typically is a function of energy and species.

ThegoalofthecalibrationistocharacterizethefunctionG($E, \theta, \phi; m$), so that from measurements of the count rates it is possible to constrain $f(E, \theta, \phi; m)$. Note that an exact inversion of the integral is rarely possible, and we can compute only the coefficients of some tailored expansion of $f(E, \theta, \phi; m)$, such as in spherical harmonics (Legendre polynomials). The accuracy of the secoefficients depends on both the raster coverage of the measurements and on the calibration.

5.1.2.3 Collimator

The EPS collimator consists of four concentric half circular plates that have holes aligned with a common point of originat the center of the EPSTOF telescope . The size and number of collimator holes define the geometric factor GF of the instrument . The many-holes collimator design minimizes the scattering of ions and electrons at the collimator while restricting the field of view (FOV) of the instrument .

GEANT4 simulation shows that the geometry factor for the total SSD are at obe 0.016 cm 2 sr. The simulation accounted for gaps between the detectors, but did not allow for the guardring dead are a between the large and small pixels or the losses in the two grids used to mount the thin foils . Hence, before gridlosses, the total large pixel geometry factor is therefore GFSSD=0.0152 cm 2 sr, and the small pixels would be 0.0008 cm 2 sr. The gridlosses are actual blockages, so these should be ...

 $included in the geometry factor \quad . \ EPS used 40-lines-per-inchgrids on the foilst hat are 86\% transmissive. Therefore, for the 12 large pixels, we have a total geometry factor of 0.862 x 0.0152, or 0.011 cm2 sr, and each large pixel is 1/12 of that, or GFSSD-Large=0.001 cm2 sr . For the 12 small pixels, we have a total geometry factor of 5.6 x 10-4 cm2 sr, or GFSSD-Small=4.7 x 10-5 cm2 sr per pixel. The current simulation does not model the scattering of lowener gyion and electron in the collimator; hence the current value of GFSSD is constant with energy and look direction . The instrument teammay revise the value of GFSSD at a later time when we develop a further under standing of the instrument response as a function of direction, energy and species.$

 $GFTOF for the Low Energy Ions (TOF-only) is roughly twice the SSD values, or $\sim 0.03 \, {\rm cm} 2 \, {\rm skj} which depends, in} \\ that the needed Transfer Factor G depends also on the counting efficiency skj which depends, in turn, on species and instrument mode. However, these values were never conclusively determined During the time of instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's TOF section suffered a failure; subsequently, EPS lost its ability to measure ions by elemental mass species and instrument check out shortly after launch, EPS's and EPS's an$

$$N(E) = \delta t * f(E) * GFSSD$$

GFSSD is now the geometric factor and represents a measure of the efficiency of the system (count rate/flux), and is constant with look direction, energy and species. This is the standard approximate conversion of countrate to intensity assuming the channel efficiency is part of the geometry factor. The shape of the energy spectrum also affects the response.

5.2 Data Product Overview

TheDDRdataproductsgeneratedbytheEPSandFIPSsubsystemsaredescribedinthissection . For alloftheDDRproductsthereisadetachedPDSlabelfile,whichdescribesthecontentsofonedata file. Each labelfile has the same base name as the data file it is describing, with the extension ".LBL" to denote a label file. The label file defines the start time and end of the observation, product creation time, and the structure of the ASCII tables . Each data file contains the data collected on given Earthday.

5.2.1 EPS Data Products

The EPS portion of the data archive consists of tables of EPS pitch angle values and pitch angle distribution functions pectrograms in the form of browse plots.

5.2.1.1 PitchAngleComputation

The pitch angle is defined as the instantaneous angle between the particle flow direction and the measured in-situm agnetic field, with 0 degrees being where the particle is traveling along the field and 180 degrees where the particle is traveling in the opposite direction. Both vectors (the particle flow direction and the magnetic field) are measured in the space craft frame. The particle flow direction for each of the six EPS look directions is obtained by using the MESSENGER SPICE pointing kernels (c-kernels) and the EPS frame kernel. The magnetic field measurements are taken from MESSENGER Magnetometer (MAG) Level 11-second averaged data; see the MAG instrument DDRSIS.

5.2.1.2 PitchAnglesProduct

The pitch angle value tables contain the pitch angle for each of the six look directions reported at the same cadence as the EPSLoRes Spectral rate measurements (see EPPSCDRSIS document). Each file thus has a one-to-one relationship with a corresponding LoRes spectral file. The pitch angle values are calculated using the odd-shaped channels. These hardware channels correspond to ions or electrons that are above the instrument threshold (no minally ~30 keV), and are not affected by prioritization in the instruments of tware. For a rapidly falling energy spectrum, these counts would be dominated at the lowest energies. The sevalues should be used in conjunction with the rate channels for which in tensity is computed. Constant normalization factors and background subtractions have been applied to the sechannels to off set any bias in the threshold setting. The formula used is Adjusted Value = Scaling *Raw Value + Off set. The constant scaling and off set values are given in Table 2.

5.2.1.3 PitchAngleDistributionSpectrogramBrowseProduct

The pitch angle distributions how ninthe pitch angle spectrogram browse plots is plotted using the pitch angle values and average daccordingly.

We averaged the particle measurement in 120-second bins. Pitch angle bins are 22.5 degrees wide, and run from 0 to 180 degrees. For the pitch angle distribution spectrogram, the color scale is normalized to the maximum flux for the time interval covered by the current plot (i.e., the intensity color scale varies between plots). The events are selected based on the selection criteria as outlined by Hoetal. [JGR, 2012] on selecting events that are at least 10 times above the instrument background at the lowest energy channel 36-57 keV. Note: The pitch angle distribution spectrogram product is a one-time delivery.

ChannelName	Scaling	Offset
SHAPED_COUNTSD00	1.0412691533565521	15.11127820611
SHAPED_COUNTSD01	1	0
SHAPED_COUNTSD02	0.7014738321304321	10.141375184059143

SHAPED_COUNTSD03		
SHAPED_COUNTSD04	1.338530421257019	18.39651709794998
SHAPED_COUNTSD05	1.3527332693338394	7.795897737145424
SHAPED_COUNTSD06	1.4293411076068878	27.3230662047863
SHAPED_COUNTSD0 7	0.805070623755455	3.0807372108101845
SHAPED_COUNTSD0 8	0.9233275651931763	32.62595450878143
SHAPED_COUNTSD09	0.5404207110404968	-0.39973045140504837
SHAPED_COUNTSD10	0.31776440143585205	22.942654877901077
SHAPED_COUNTSD11	0.15649390034377575	-22.306587459519506

Table 2 Constant values used to adjust the countrate for the shaped channels. D01 is the channel to which all others were fit in order to arrive at these values, while D03 is unlike the others and cannot be fit, so it remains unadjusted.

5.2.2 FIPS Data Products

Except where noted below, the sed at a products are derived from individual event (PHA) words, which have full angle and TOF information. Using PHA words enables the finest possible separation into ion species as well as distinguishing incident angles. No is eremoval in E/q – TOF space is performed before a ccumulation for all ions except E/q +, for which no is eremoval is not needed.

The orbital regions for which data is provided varies by product, as described below. Within these, data is provided for time steps whenever the instrument was operating in a nominal mode and when the data was of sufficient quality for scientificanalysis. The seconditions are met for the vast majority of times throughout the mission. In particular, many of the seproducts are provided in regions determined by hand-picked or modeled bowshock and magnetopause boundaries. Boundary picking was done by the MAG team at a level of accuracy sufficient for the relatively low time resolution FIPS data (10s-60s). Where the sewere not available, the model given by Slavinet al. [14] is used.

5.2.2.1 DifferentialEnergyFluxSpectra

Differentialenergyfluxspectrainunitsof(keV/e) $^{-1}$ sec $^{-1}$ cm $^{-2}$ sr $^{-1}$ areprovidedforall5ionspecies throughouttheentiremagnetosphereandmagnetosheathregionscrossedbythespacecraftduring eachorbit. These spectracover the fullenergy perchargerange of FIPS, which ws 0.046 -13.3 keV/eduring most of the mission. See the FIPA_EYYYYDOY. DAT file for exact energy range. They are derived in a manner completely analogous to CDR differential energy flux spectra ($dJ_{i,s}/dE_i$) except that the total number of PHA event words for a given species at each E/q step is used in place of on-board accumulated count rates (except protons). See EPPS_FIPS_EDR2CDR. DOCX for more details. For protons, this product is derived from on-board accumulated rates just as for the CDR version, so that for **protons only** this product is a duplicate of the CDR version. This duplication is made to allow users to primarily use DDR data only in their scientific analysis.

$$\frac{dJ_{i,s}}{dE} = \frac{J_{i,s}}{(E/q)_i}$$

5.2.2.2 Observed Density

Observed density ($n_{\text{obs,s}}$, in cm⁻³) for all ion species is provided for all 5 ion species throughout all orbit regions. This product serves to determine the spatial and temporal distribution of these ions and is the foundation DDR data product for FIPS. For protons, this product is derived from onboard accumulated rates rather than PHA event words, for much higher signal to noise ratio.

Observed densities are computed from differential energy flux spectra $J_{i,s}$ in two stages: First, $dJ_{i,s}/dE_i$ is converted to phase-space density ($f_{i,s}$, s^3 m⁻⁶)

$$f_{i,s} = 6.2414 \times 10^{19} \left(\frac{m_s}{v_{i,s}^2} \right) \left(\frac{dJ_{i,s}}{dE_i} \right)$$

Second, $f_{i,s}$ integrated over all velocities ($v_{i,s}$) and solid angle ($\Delta\Omega$) for a full scan to yield observed density ($n_{obs,s}$) for species s

$$n_{obs,s} = \sum_{i} f_{i,s} v_{i,x}^{2} (Dv)_{i,s} (DW)$$

It is important to note that no correction has been made for limited field of view in this calculation. This correction requires knowledge of the true velocity distribution function in which the measured particles reside and is therefore beyond the scope of this data product. The net result is that nobs values reported are not in general equal to the ambient plasma density. More information about calculation of n_{obs} and the limited FIPS field of view can be found in [10, 11].

5.2.2.3 Angular Flux Maps

Integrated ion flux, $J_s(\theta,\phi)$, as a function of flow direction in MSO coordinates are provided for all 5 ion species throughout the entire magnetosphere and magnetosheath regions crossed by the spacecraft during each orbit. This product replaces the Arrival Direction Histogram (ARRDIR) product as it more a more useful representation of plasma behavior. Angular Flux Maps may be used in conjunction with pitch angle distributions to understand the location and motion of ions

 $relative to Mercury. The discrete integration of phase spaced ensity (\begin{tabular}{l} f) is performed over ion speed \\ (v) in the standard way, leaving the two angular dimensions unchanged: \end{tabular}$

$$J_{s}(\theta,\phi) = \sum_{i} f_{i,s}(\theta,\phi) v_{i,s}^{3} dv_{i}$$

Very low count rates can make this product difficult to interpret. For some ions, this situation is mitigated by summing over several, of ten many, FIPS scan suntil sufficient counts are available. In other cases, counts for a particular ionare so low the required number of counts would require summation over large regions of space, making the arrival directions again difficult to interpret. Therefore, these cases are not included in the data.

Summingovermultiplescansbringswithittheneedtonormalizethearrivaldirection distributions for the amount of time that aparticular incident angle could be observed. Innormal operations, the MESSENGER spacecraft rotates around the spacecraft Y-axis ($Y_{\rm MSGR}$) to optimize viewing for different instruments. FIPS or ientation is fixed relative to the spacecraft, so that the FIPS FOV rotates with the MESSENGER spacecraft. This results in significant variation in observing time for look directions in MSO coordinates. These variations in observing time have been normalized out of the fluxes included in this data product. Despite this normalization, the finite sensitivity of FIPS can still result in reduced measured ion flux from arrival directions with low observing time. In practice, this does not usually affect the typically qualitative interpretation of these maps. The user could mitigate this effect by reducing the variability of observing time: Only flux from directions that are viewed within a fixed fraction (e.g. 0.1) of the maximum observation time for a given time accumulation could be included in arrival direction his to grams. This restriction is not applied to delivered data from this data product.

The FLUXMAP data product is constructed and normaliz edas follows:

- 1. Two arrays representing a full 4π steradian FOV are created at the 10° native angle binning ofFIPS.Thisresultsinan18x36elementarrayforthepolarandazimuthalanglebins, respectively.Thefirstoftheseisusedfortheaverageflux,FLUX,whilethesecondisused fortheaveragemeasurementtime,VIEWTIME.
 - a. TheFLUXmatrixforeachenergyscaniscreatedfromPHAsasfollows:

- i. Findtherotationmatrix,ROTMSO,(describedbelow)fortheday (YYYYDOY)andindexwhichcorrespondstothescan
- ii. Formtheunitvelocityvector:
 - Converttheincidentpolarandazimuthalanglesoftheeventfrom FIPSsphericalcoordinatestoFIPSCartesiancoordinates.Usethe usualmannerforsphericaltoCartesianconversion,using1forther component.
 - 2. Invertbymultiplyingallcomponentsby-1.ThischangestheFOV coordinatestoflowdirection.
- iii. RotatetoMSOviamultiplicationbytheROTMSO.
- iv. ConverttheresultantMSOvectortosphericalcoordinates.Roundangles downtothenearest10°.
- v. AddthefluxvaluemultipliedbythePHAsolidangleintoatemporary18x 36matrixusingtherounded-downangles.
- vi. ConvertentireFIPSsphericalMCPmaptoMSOusingthestepsfrom(ii)for each coordinate pair. Roundangles down to the nearest 10°.
- $vii. \ \ Add the MCP pixel solid angles into a second temporary 18x36 matrix using the rounded-down angles.$
- viii. Dividethetemporaryfluxbythetemporarysolidangle,,element- by-element.
- b. TheVIEWTIMEmatrixforeachenergyscaniscreatedasfollows:
 - i. Calculatetheaveragestepmeasurementtimeforthescan
 - ii. AddthatvalueintoeveryelementoftheVIEWTIMEmatrix
 - $iii. \ \ Setto 0 any matrix location for which the corresponding solid angle matrix \\ value is 0 (this will remove unobserved locations from the result).$
- 2. SumFLUXandVIEWTIMEoverallaccumulationscans
- 3. Formtheproperlynormalized product by dividing (element by element) the FLUX (summed overscans) and VIEWTIME (summed overscans).

$$FLUXMAP = \frac{\sum FLUX}{\sum VIEWTIME}$$

5.2.2.4 ArrivalDirection(RetiredProduct)

This product class was retired after the Mercury Orbit Year 2 mission phase. The Angular Flux Map products described in the previous section supersede it. Since the Arrival Direction products remain inversion V1.0 of the dataset in the FIPSDDR archive volume at the PDS, a description of the misretained in this document.

Thisproductprovides observed density as a function of arrival direction in the instrument frame (incident polar and azimuthal angle) in MSO coordinates for selected ions and regions around Mercury. The discrete integration of phases pacedensity (f) is performed in the velocity dimension (v) only, leaving the two angular dimensions unchanged:

$$n_{obs,s}(\theta,\phi) = \sum_{i} f_{i,s}(\theta,\phi) v_{i,s}^{2} dv_{i}$$

Theabove quantity, termed "observed density",isanalogoustodensityineachangularbinformed onlyfromthecountsobservedinthatbin.ItdiffersfromthatperformedintheAngularFluxMap productbyonepowerof vwhichintroducedinthatproducttochangefromadensitytoaflux quantity.

The Arrival Direction Histogram product is identical in several ways to the Angular Flux Maps product that replaced it: It can be difficult to interpret due to very low count rates, a problem which is partially mitigated by summing over FIPS scans. This summing then requires proper normalization, performed in exactly the same manner as described in detail for Angular Flux Maps. This product may be used in conjunction with pitch angle distribution stounderstand the location and motion of ions relative to Mercury.

The ARRDIR data product is constructed and normaliz edex actly as described above for the Angular Flux Maps, with two exceptions (with the ARRDIR array replacing the FLUX MAP array):

1. ThevelocityvectormustNOTbeinvertedbyreversingthesignofeachcomponent.

2. MultiplytheweightfromeachPHAisaddedintotheARRDIRarray.Itis **not**multipliedby theparticlespeed. (Modifystep2.b.iv.above.)

5.2.2.5 *PitchAngleDistributions*

For anioninthepresenceofamagnetic field, the angle between the velocity vector and the local magnetic field direction is referred to a spitch angle . For populations of plasmaions, pitch angle distribution histograms can be for med by counting the number of ions within a given pitch angle range (PCHANG data product). This histogram may also be separated in to measured E/q bins, for mingenergy-resolved pitch angle distributions (ERPCHANG data product) . Pitch angle distributions give information on the velocity of ions along magnetic field lines, character of velocity distributions and the general plasma environment. Care must be taken when interpreting pitch angle distributions in the instrument frame (such as the sellow when the plasma has a nonnegligible bulk velocity. Both the PCHANG and ERPCHANG products are provided for all 5 ion species throughout the entire magnetos phereand magnetos heath regions crossed by the space craft during each or bit.

 $\label{thm:constraint} The pitch angle distribution for a given time period consists of a histogram of these angles in bins for all the ionevents (PHA words) in the time period. Energy-resolved pitch angle distributions are also separated into the native FIPS <math>E/q$ stepping bins. Pitch angle distributions are provided for selected ions and time periods, when sufficient statistics exist to make the product meaning fulfor scientific studies.

The PCHANG data production or maliz edas follows:

- $1. \label{lem:condition} Two arrays representing pitch angles 0-180° are created at 10° binning, resulting in two 18-elementarrays. The first of these is used for the average flux, FLUX, while the second is used for viewing normalization, VIEW TIME.$
 - $a. \quad The FLUX array for each energy scan is created from PHAs as follows:$
 - i. Findtherotationmatrix,ROTMSO,(describedbelow)fortheday (YYYYDOY)andindexwhichcorrespondstothescan
 - ii. Formtheunitvelocityvector:
 - $1. \quad Convert the incident polar and a zimuthal angles of the event from FIPS spherical coordinates to FIPS Cartesian coordinates. Use the$

usual manner for spherical to Cartesian conversion, using 1 for the recomponent.

- 2. Invertbymultiplyingallcomponentsby-1.ThischangestheFOV coordinatestoflowdirection.
- iii. RotatetoMSOviamultiplicationbytheROTMSO.
- $iv. \ \ Average the MAGCDR vectors (already in MSO) that fall within the FIPS scantimeto one value$
- v. CalculatetheangularseparationbetweenthisvectorandtheaverageMAG vector.Rounddownangletonearest10°.
- vi. AddthefluxvaluemultipliedbythePHAsolidangleintoatemporary18 elementarrayusingtherounded-downangles.
- vii. ConvertentireFIPSsphericalMCPmaptoMSOusingthestepsfrom(ii)for each coordinate pair. Roundangles down to the nearest 10°.
- viii. CalculatetheangularseparationbetweentheMCPvectorsandtheaverage MAGvector.Rounddownangletonearest10°.
- ix. AddtheMCPpixelsolidanglesintoasecondtemporary18elementarray usingtherounded-downangles.
- x. Dividethefluxtempbythesolidangletemp, element- by-element.
- b. TheVIEWTIMEarrayforeachenergyscaniscreatedasfollows:
 - i. Calculatetheaveragestepmeasurementtimeforthescan
 - ii. AddthatvalueintoeveryelementoftheVIEWTIMEarray.
 - iii. Setto0anyarraylocationforwhichthecorrespondingsolidanglearray valueis0(thiswillremoveunobservedlocationsfromtheresult).
- 2. SumFLUXandVIEWTIMEoverallaccumulationscans
- $3. \ \ Formthe properly normalized product by dividing (element by element) the FLUX (summed overs cans) and VIEWTIME (summed overs cans).$

$$PCHANG = \frac{\sum FLUX}{\sum VIEWTIME}$$

The ERPCHANG product is formed in an analogous fashion, with the additional separation by E/q step. In this case, the weighting value in the 'FLUX' numerator is average phase space density.

5.2.2.6 Kinetic Properties

Under several conditions, full number density, n, temperature, T, and pressure, P, can be calculated directly from counts. (In contrast, when these conditions are not fulfilled, only the partial, observed densities, described above in 5.2.2.2, can be determined.) The conditions are as follows:

- 1) There must be sufficient counts to produce a well-defined energy spectrum. When necessary, counts from multiple scans are summed to meet this criterion.
- 2) The plasma must be subsonic. That is,

$$\frac{v_b}{v_{th}} < 0.5$$

where v_b is the plasma bulk velocity and V_{th} is the plasma thermal velocity.

3) The plasma must be nearly isotropic. That is,

$$0.5 < \frac{T_{1}}{T_{||}} < 5$$

where T_{\perp} is the plasma temperature perpendicular to the magnetic field and $T_{||}$ is the plasma temperature parallel to the magnetic field.

This data is provided where these assumptions are most likely to hold: a) throughout the magnetosphere and b) the dayside magnetosheath within a 45° cone angle around the $X_{\rm MSO}$ axis. Furthermore, the data are averaged over multiple FIPS scans so as to minimize the effect on the recovered n and T of transients (e.g. high speed flows) which violate the assumptions. **However**, for studies involving features near the time scale of the averaged data and/or inside the magnetosheath, **users are strongly cautioned** to evaluate the validity of the assumptions themselves. This is most easily done by comparing velocity distribution functions, computed from differential energy flux spectra and averaged to the same period as this data, to those formed from a non-drifting Maxwell-Boltzmann velocity distribution formed from the recovered n and T (with bulk speed of zero).

For protons, this product is derived from on-board accumulated rates rather than PHA event words, for much higher signal to noise ratio. Data is summed over multiple scans to increase signal to noise. Furthermore, this data is produced only those accumulations which exceed a minimum

countlevel (20 counts). Extensive testing has shown that computing these products from PHA event words for ions heavier than protons is not practical at fixed timesteps of the order of a few minutes. Therefore, these ions are not included in this product.

When these conditions are fulfilled, we compute n and v_{th} directly from measured velocity distribution functions (formed from differential energy flux spectra) using a numerical method of solving the system of momente quations [Gershman et al. , 2013]. Tand Parethen calculated using the usual relations:

$$T = m v_{th}^2 / k_b$$
$$P = n k_b T$$

where k_b is the Boltzmann constant. The units of the recovered plasma parameters are cm and nPaforn, T, and P, respectively. The quoted uncertainties of these parameters are a function of only number of counts used to create each distribution, following Gershman et al., 2013. There are small additional uncertainties that result as the measured energy spectra approach the limits of conditions (2) and (3) above. The sead ditional uncertainties are difficult to quantify and not included in the reported uncertainties.

5.2.2.7 ViewingNormalization

The FIPS Viewing Normalization data product contains a rotation matrix (ROTMSO) from FIPS cartesiant om SOcoordinates, for each FIPS energy scan. This matrix can be used to rotate ion incident angles in CDRPHA data into MSO coordinates needed for producing normalized directional maps (e.g. FLUXMAP or PCHANG) for arbitrary time resolutions, in multiples of 10s.

5.3 Data Processing

5.3.1 Data Processing Level

There isoneEPPSPDSDocumentationArchiveVolumeandoneEPPSPDSDataArchiveVolume.

Thedatavolumecontainslevel4CODMACdataproducts,alsoknownasDDRs.Eachproducthas a uniquefilenameandconformstothefilenamingconventioninsection 6.5.AllDDRproductswere storedattheMESSENGERScienceOperationsCenter(SOC)duringtheMESSENGERmission .

VolumesweretransferredtothePDSPPINodefollowingtheprocedureinsection 5.3.3.

5.3.2 Data Product Generation

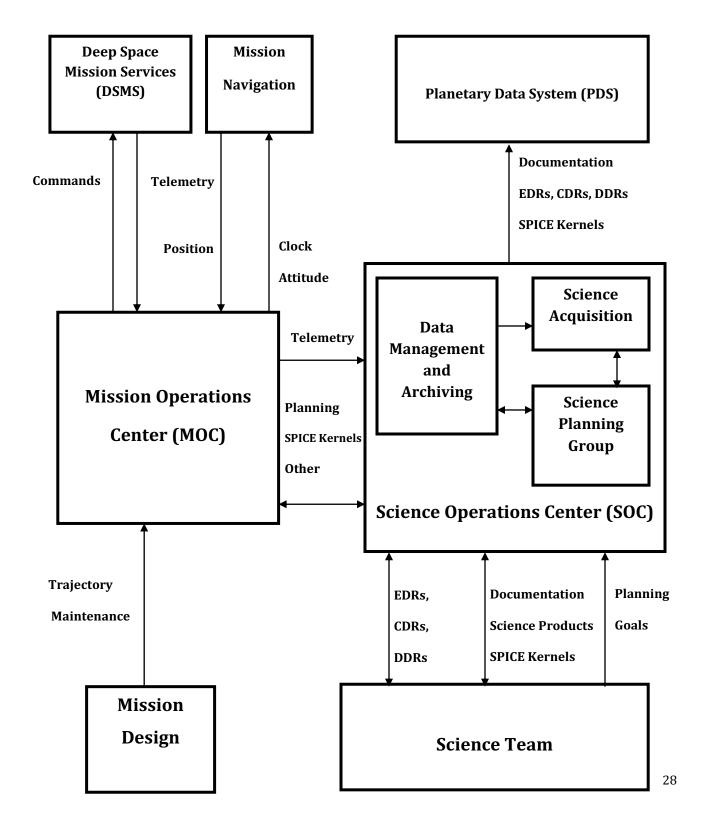
The EPPSDDR files were produced by the EPS and FIPS teams. A Java program derived from the MIDL (Mission Independent Data Layer) analysis software developed by APL was used to generate the DDRs. The FIPS data were produced using Interactive Data Language (IDL) software routines developed at the University of Michigan. The DDR data products were made available to the MESSENGERS cience Teamfor initial evaluation and validation. At the end of the evaluation and the transfer of the transfer

validationperiod,thedatawereorganizedandstoredinthedirectorystructuredescribedin section 6.8fortransfertothePPINode.Thetransmittalprocessisdescribedinsection, 5.3.3. An initialreleaseofthedocumentationvolumeaccompaniedtheinitialreleaseofthedatavolume. Thereafter,updatestothedocumentationvolumeweremadewitheachdatadeliverytodocument thedataqualityforthedelivery,changestoproductsincludingcalibrationupdates,andother updatesasappropriate.PDSprovidespublicaccesstothedataproductsthroughitsonline distributionsystem.Theseproductssupportengineeringanalysis,directscienceanalysis,and constructionofotherscienceproducts.

5.3.3 Data Flow

The MESSENGERSO Coperates under the auspices of the MESSENGER Project Scient is stop landata acquisition, generate, and validate data archives. The SOC supports and works with the Mission Operations Center (MOC), the Science Team, instruments cientists, and the PDS.

Figure 1 MESSENGER data flow shows the flow of data within the MESSENGER project and out to PDS. The MOChandles raw data flow to and from the MESSENGER space craft and the SOC converts the raw telemetry into EDRs, which are subsequently converted into CDRs and DDRs by the Science Team. Documentation, CDRs, and DDRs are delivered to the PDS Planetary Plasma Interactions (PPI) node. All SPICE kernels used in CDR and DDR processing are delivered to the PDS Navigation and Ancillary Information (NAIF) node. The delivery process is detailed below.



The MESSENGERSOC delivered data for the EPPSDDR data volume to the PDSPPIN ode in standard product packages. Each package comprises data and ancillary data files, organized into directory structures consistent with the volume design described in section 6.8. The initial release contained the documents and required files for the EPPS documentation volume, organized into directory structures as described in section 6.7. Subsequent releases to the EPPS documentation volume contained updates as appropriate.

In preparation for delivery, the directory structure is compressed into a single "zip archive" file for transfertothePDSnode. The ziparchive preserves the directory structure internally so that it can be recreated after delivery to the PDS node. Also included in the transfer is a check sumfile created using the MD5 algorithm. This provides an independent method of verifying the integrity of the zip file after it has been sent. Within days of receipt of the delivery the PDS node acknowledges receipt of the archive and check sumfile. If a cknowledge ment is not received, or if problems are reported, the MESSENGERSOC immediately takes corrective action to affect successful transfer. Delivery size determines the transfer mechanism: electronic or shipping a hard drive.

The PDS node uncompresses the ziparchive file and checks for data integrity using the check sum file. The node performs any additional verification and validation of the data provided and reports any discrepancies or problems to the MESSENGERSOC. The node performs the sechecks within about two weeks from receipt of the delivery. After inspection has been completed to the satisfaction of the PDS node, the node is sue san acknowledgement of successful receipt of the data to the MESSENGERSOC.

Following receipt of a data delivery the PDS node organizes the data into a PDS volume archive structure within its online data system. Newly delivered data are made available publicly from PDS once accompanying labels and other documentation have been validated.

5.3.4 Labeling and Identification

 $The PDS label conforms to PDS version 3.8 standards. For more information about this standard consult the PDS standards Reference Document. The label is detached and in a separate PDS label file. The purpose of the PDS label is to describe the data product and provide ancillary information about the data product. There is a PDS label file for every EPPSDDR data file. There is one DATA_SET_ID assigned to the EPPSDDR data. The DDRs are further grouped into data products and are identified by the STANDARD_DATA_PRODUCT_ID keyword and the file naming convention, section 6.5. Example label file content is shown here for every DDR data product. Note that the data are contained within an ASCII table and the details of the table structure are described by an external ASCII format file (*.FMT). The columns in each format file are described separately in the Appendix.$

5.3.4.1 EPSPitchAnglesLabel

AsampleEPSPitchAnglesDDRfilelabelisshownbelow:

```
PDS VERSION ID = "PDS3"
/* ** FILE FORMAT ** */
           = 5798
FILE RECORDS
RECORD_TYPE
                  = FIXED_LENGTH
RECORD BYTES
                   = 167
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
                   = "EPSP A2012010DDR V1"
PRODUCT_ID
PRODUCT VERSION ID = "V1"
PRODUCT_CREATION_TIME = 2012-05-09T21:04:27
PRODUCT_TYPE = "DDR"
STANDARD DATA PRODUCT ID = "EPS PITCH ANGLES DDR"
SOFTWARE_NAME = "MIDLMessengerDDRGenerator"
SOFTWARE_VERSION_ID = "1.0"
INSTRUMENT NAME = "ENERGETIC PARTICLE SPECTROMETER"
INSTRUMENT_ID = "EPS"
DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-EPS-DDR-V1.0"
DATA_SET_NAME = "MESSENGER E/V/H/SW EPPS CALIBRATED EPS DDR
V1.0"
MISSION_PHASE_NAME
                  = "MERCURY ORBIT"
                   = "MERCURY"
TARGET NAME
START TIME
                   = 2012-010T00:00:49
                   = 2012-010T23:59:45
STOP TIME
```

SPACECRAFT_CLOCK_START_COUNT = "234641115"

SPACECRAFT_CLOCK_STOP_COUNT = "234727451"

^HEADER = ("EPSP A2012010DDR V1.TAB", 1)

^ASCII_TABLE = ("EPSP_A2012010DDR_V1.TAB", 2)

OBJECT = HEADER

HEADER_TYPE = TEXT

INTERCHANGE FORMAT = "ASCII"

RECORDS = 1

BYTES = 167

DESCRIPTION = "The first record of this

file is the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII_TABLE

COLUMNS = 7

ROW_BYTES = 167

ROWS = 5798

DESCRIPTION = "

This table contains Pitch Angles between the measured flow vector direction and the magnetic field for each of the 6 sectors of the MESSENGER EPS instrument. The complete column definitions are contained in an external file found in the LABEL directory of the archive volume. Additional details are contained in the DDR SIS document."

NOTE = "Data Quality: 0"

```
^STRUCTURE = "EPS_PITCH_ANGLES.FMT"

END OBJECT = ASCII TABLE
```

END

5.3.4.2 EPSPitchAngleSpectrogramLabel

A sample EPSP itch Angle Spectrogram DDR file label is shown below:

```
PDS VERSION ID = "PDS3"
/* ** FILE FORMAT ** */
                        = UNDEFINED
RECORD TYPE
INTERCHANGE FORMAT = BINARY
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
PRODUCT_ID = "EPS_PAS_2012074205045_V1"
PRODUCT_VERSION_ID = "V1"
PRODUCT_CREATION_TIME = 2012-05-09T17:00:00
PRODUCT_TYPE = "BROWSE"
STANDARD_DATA_PRODUCT_ID = "EPS_PITCH_ANGLE_SPECTROGRAM_DDR"
SOFTWARE NAME = "MIDLMessengerDDRGenerator"
SOFTWARE_VERSION_ID = "1.0"
INSTRUMENT HOST NAME = "MESSENGER"
INSTRUMENT_NAME = "ENERGETIC PARTICLE SPECTROMETER"
INSTRUMENT_ID
                    = "EPS"
DATA_SET_ID
                     = "MESS-E/V/H/SW-EPPS-3-EPS-DDR-V1.0"
DATA SET NAME
                     = "MESSENGER E/V/H/SW EPPS CALIBRATED EPS DDR V1.0"
MISSION PHASE NAME
                     = "MERCURY ORBIT"
                     = "MERCURY"
TARGET NAME
```

```
START TIME
                   = 2012-03-14T20:50:45
      STOP TIME
                             = 2012-03-15T00:23:45
      SPACECRAFT CLOCK START COUNT = "240245710"
      SPACECRAFT CLOCK STOP COUNT = "240258490"
      ^DOCUMENT
                           = "EPS_PAS_2012074205045_V1.PNG"
      OBJECT
                           = DOCUMENT
                         = "EPS_PAS_2012074205045 V1"
        DOCUMENT NAME
        DOCUMENT FORMAT
                          = PNG
        DOCUMENT TOPIC TYPE = "BROWSE IMAGE"
        INTERCHANGE FORMAT = BINARY
        PUBLICATION DATE = 2012-05-09T17:00:00
        SOURCE PRODUCT ID = {
            "EPSL R2012074EDR V1.DAT",
            "MAGSC SCIAVG12075 01 V00.TAB",
            "MAGSC SCIAVG12074 01 V00.TAB",
            "EPSL R2012075EDR V1.DAT"
        LINES
                 = 400
       LINE SAMPLES = 850
                      = MSB UNSIGNED INTEGER
        SAMPLE TYPE
        SAMPLE BITS
                         = 8
       DESCRIPTION = "PNG file containing a spectrogram representation of
the summation of the shaped count rates in detectors 0-10 except detector 3. The
counts are binned in time (120 s) and pitch angle (22.5 deg). Constant normalization
factors and background subtractions (see SIS for Table) have been applied to these
rates."
      END OBJECT = DOCUMENT
      END
```

5.3.4.3 FIPSPitchAnglesLabel(PCHANG)

AsampleFIPSPitchAnglesDDRfilelabelisshownbelow:

```
= "PDS3"
     PDS VERSION ID
     /* ** FILE FORMAT ** */
     FILE RECORDS
                          = 1350
     RECORD TYPE
                          = FIXED_LENGTH
                          = 2799
     RECORD BYTES
     /* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
     PRODUCT ID
                          = "FIPS PCHANG 2012001 DDR V01"
     PRODUCT VERSION ID = "V1"
     PRODUCT_CREATION_TIME = 2012-05-08T23:35:42
     PRODUCT_TYPE = "DDR"
     STANDARD_DATA_PRODUCT_ID = "FIPS_PCHANG_DDR"
     SOFTWARE NAME = "mfips decode pha.pro"
     SOFTWARE_VERSION_ID = "1.0"
     INSTRUMENT HOST NAME = "MESSENGER"
     INSTRUMENT_NAME = "FAST IMAGING PLASMA SPECTROMETER"
     DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
     DATA_SET NAME
                           = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
     SOURCE PRODUCT ID
                          = "FileByFile"
    MISSION_PHASE_NAME
                          = "MERCURY ORBIT"
                           = "MERCURY"
     TARGET NAME
     START TIME
                         = 2012-01-01T00:00:00
```

STOP TIME = 2012-01-01T23:59:59

SPACECRAFT_CLOCK_START_COUNT = "233863466"

SPACECRAFT_CLOCK_STOP_COUNT = "233949802"

^HEADER = ("FIPS_PCHANG_2012001_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_PCHANG_2012001_DDR_V01.TAB", 4)

OBJECT = HEADER

HEADER_TYPE = TEXT

INTERCHANGE_FORMAT = "ASCII"

RECORDS = 3

BYTES = 8397

DESCRIPTION = "The first 3 records of this

file are the header section. The header contains column

headings to improve usability."

END_OBJECT = HEADER

OBJECT = ASCII_TABLE

COLUMNS = 3

INTERCHANGE FORMAT = ASCII

ROW BYTES = 2799

ROWS = 1350

DESCRIPTION = "

This table contains FIPS Flux-Pitch angle histograms."

^STRUCTURE = "FIPS_PCHANG_DDR.FMT"

END_OBJECT = ASCII_TABLE

END

5.3.4.4 Energy-ResolvedPitchAngleDistributions(ERPCHANG)Label

AsampleFIPSERPCHANGDDRfilelabelisshownbelow:

```
PDS VERSION ID
            = "PDS3"
/* ** FILE FORMAT ** */
FILE_RECORDS = 5
RECORD TYPE = FIXED LENGTH
RECORD BYTES = 16218
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
PRODUCT_ID
                    = "FIPS_ERPCHANG_2011174_V1"
PRODUCT_VERSION_ID = "1"
PRODUCT_CREATION_TIME = 2014-11-06T16:00:00
                     = "DDR"
PRODUCT TYPE
STANDARD_DATA_PRODUCT_ID = "FIPS_ERPCHANG"
SOFTWARE NAME
                     = "MFIPS DDR SAMPLE.PRO"
SOFTWARE_VERSION_ID = "0.1"
INSTRUMENT_HOST_NAME = "MESSENGER"
INSTRUMENT_NAME = "ENERGETIC PARTICLE AND PLASMA SPECTROMETER"
DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
DATA_SET_NAME
                      = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
MISSION_PHASE_NAME = "MERCURY ORBIT"
          = "MERCURY"
TARGET NAME
START TIME = 2011-06-23T10:45:40.419666
        = 2011-06-23T22:53:43.420605
STOP TIME
SPACECRAFT CLOCK START COUNT = "1/217313408.800"
```

SPACECRAFT_CLOCK_STOP_COUNT = "1/217357091.000"

^HEADER = ("FIPS_ERPCHANG_2011174_DDR_V01.TAB", 1)

^ASCII TABLE = ("FIPS ERPCHANG 2011174 DDR V01.TAB", 4)

OBJECT = HEADER

HEADER_TYPE = TEXT

INTERCHANGE_FORMAT = "ASCII"

RECORDS = 3

BYTES = 48654

DESCRIPTION = "The first four records of this

file are the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII_TABLE

COLUMNS = 7

INTERCHANGE FORMAT = ASCII

ROW_BYTES = 16218

ROWS = 2

DESCRIPTION = "

This table contains 2 dimensional pitch angle histograms as

described in EPPS DDR SIS."

^STRUCTURE = "FIPS_ERPCHANG_DDR.FMT"

END_OBJECT = ASCII_TABLE

END

5.3.4.5 FIPSEnergySpectraLabel

A sample FIPS Energy Spectra DDR file label is shown below:

```
= "PDS3"
    PDS VERSION ID
    /* ** FILE FORMAT ** */
    FILE RECORDS
                         = 1350
    RECORD TYPE
                         = FIXED_LENGTH
                         = 4824
    RECORD BYTES
    /* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
    PRODUCT ID
                         = "FIPS ESPEC 2012001 DDR V01"
    PRODUCT VERSION ID = "V1"
    PRODUCT_CREATION_TIME = 2012-05-08T23:35:41
    PRODUCT_TYPE = "DDR"
    STANDARD_DATA_PRODUCT_ID = "FIPS_ESPEC_DDR"
    SOFTWARE NAME = "mfips decode pha.pro"
    SOFTWARE VERSION ID = "1.0"
    INSTRUMENT_NAME = "FAST IMAGING PLASMA SPECTROMETER"
    DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
    DATA_SET NAME
                         = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
    SOURCE PRODUCT ID
                   = "FileByFile"
    MISSION_PHASE_NAME
                         = "MERCURY ORBIT"
                         = "MERCURY"
    TARGET NAME
    START TIME
                      = 2012-01-01T00:00:00
```

STOP TIME = 2012-01-01T23:59:59

SPACECRAFT_CLOCK_START_COUNT = "233863466"

SPACECRAFT_CLOCK_STOP_COUNT = "233949802"

^HEADER = ("FIPS_ESPEC_2012001_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_ESPEC_2012001_DDR_V01.TAB", 4)

OBJECT = HEADER

HEADER_TYPE = TEXT

RECORDS = 3

BYTES = 19094472

DESCRIPTION = "The first 3 records of this

file are the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII_TABLE

COLUMNS = 7

INTERCHANGE FORMAT = ASCII

ROW BYTES = 4824

ROWS = 1350

DESCRIPTION = "

This table contains FIPS energy spectra for selected ion species."

^STRUCTURE = "FIPS_ESPEC_DDR.FMT"

END_OBJECT = ASCII_TABLE

END

5.3.4.6 FIPSObservedDensityLabel

A sample FIPSObserved Density DDR file label is shown below:

```
PDS VERSION ID
                = "PDS3"
    /* ** FILE FORMAT ** */
    FILE RECORDS
                        = 1350
                 = FIXED_LENGTH
    RECORD TYPE
                        = 216
    RECORD BYTES
    /* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
    PRODUCT ID
                        = "FIPS NOBS 2012001 DDR V01"
    PRODUCT VERSION ID = "V1"
    PRODUCT_CREATION_TIME = 2012-05-08T23:35:42
    PRODUCT_TYPE = "DDR"
    STANDARD_DATA_PRODUCT_ID = "FIPS_NOBS_DDR"
    SOFTWARE_NAME = "mfips_decode_pha.pro"
    SOFTWARE VERSION ID = "1.0"
    INSTRUMENT_NAME = "FAST IMAGING PLASMA SPECTROMETER"
    DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
    DATA_SET NAME
                         = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
    SOURCE PRODUCT ID
                   = "FileByFile"
    MISSION_PHASE_NAME
                         = "MERCURY ORBIT"
                         = "MERCURY"
    TARGET NAME
    START TIME
                      = 2012-01-01T00:00:00
```

STOP TIME = 2012-01-01T23:59:59

SPACECRAFT_CLOCK_START_COUNT = "233863466"

SPACECRAFT_CLOCK_STOP_COUNT = "233949802"

^HEADER = ("FIPS_NOBS_2012001_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_NOBS_2012001_DDR_V01.TAB", 4)

OBJECT = HEADER

HEADER_TYPE = TEXT

INTERCHANGE FORMAT = "ASCII"

RECORDS = 3

BYTES = 648

DESCRIPTION = "The first 3 records of this

file are the header section. The header contains column

headings to improve usability."

END_OBJECT = HEADER

OBJECT = ASCII_TABLE

COLUMNS = 20

INTERCHANGE FORMAT = ASCII

ROW BYTES = 216

ROWS = 1350

DESCRIPTION = "

This table contains FIPS differential energy intensities for selected ion species."

^STRUCTURE = "FIPS NOBS DDR.FMT"

END OBJECT = ASCII TABLE

END

5.3.4.7 FIPSArrivalDirectionLabel(RetiredProduct)

A sample FIPS Arrival Direction DDR file label is shown below:

```
PDS VERSION ID
              = "PDS3"
/* ** FILE FORMAT ** */
FILE RECORDS
                      = 42
                   = FIXED_LENGTH
RECORD TYPE
RECORD BYTES
                      = 9158
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
PRODUCT_ID
                     = "FIPS_ARRDIR_2012054 DDR V01"
PRODUCT_VERSION_ID = "01"
PRODUCT_CREATION_TIME = 2013-06-04T21:22:22
PRODUCT TYPE
          = "DDR"
STANDARD_DATA_PRODUCT_ID = "FIPS_ARRDIR_DDR"
SOFTWARE NAME = "mfips ddr sample.pro"
SOFTWARE_VERSION_ID = "0.1"
INSTRUMENT HOST NAME = "MESSENGER"
INSTRUMENT_NAME = "FAST IMAGING PLASMA SPECTROMETER"
             = "FIPS"
INSTRUMENT_ID
              = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V1.0"
DATA_SET_ID
DATA_SET_NAME
                       = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V1.0"
SOURCE PRODUCT ID
                      = "FileByFile"
MISSION_PHASE_NAME
                      = "MERCURY ORBIT"
                      = "MERCURY"
TARGET NAME
```

SPACECRAFT CLOCK START COUNT = "1/238523075.000"

START TIME = 2012-02-23T22:20:08.845

SPACECRAFT_CLOCK_STOP_COUNT = "1/238524872.000"

STOP_TIME = 2012-02-23T22:50:05.845

^HEADER = ("FIPS_ARRDIR_2012054_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_ARRDIR_2012054_DDR_V01.TAB", 4)

OBJECT = HEADER

HEADER TYPE = TEXT

RECORDS = 3

BYTES = 9158

DESCRIPTION = "The first three records of this

file are the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII TABLE

COLUMNS = 7

ROW_BYTES = 9158

ROWS = 39

DESCRIPTION = '

This table contains FIPS ion flux as a function of arrival direction which have been accumulated in time enough to be meaningfully interpreted."

^STRUCTURE = "FIPS ARRDIR DDR.FMT"

```
END_OBJECT = ASCII_TABLE
```

END

5.3.4.8 FIPSAngularFluxMapLabel

AsampleFIPSAngularFluxMapDDRfilelabelisshownbelow:

```
= "PDS3"
PDS VERSION ID
/* ** FILE FORMAT ** */
FILE_RECORDS
          = FIXED_LENGTH
RECORD_TYPE
              = 9162
RECORD_BYTES
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
PRODUCT_ID
                     = "FIPS FLUXMAP 2011174 V1"
PRODUCT_VERSION_ID = "1"
PRODUCT_CREATION_TIME = 2014-11-06T16:00:00
                    = "DDR"
PRODUCT TYPE
STANDARD_DATA_PRODUCT_ID = "FIPS_FLUXMAP"
SOFTWARE_NAME = "MFIPS_DDR_SAMPLE.PRO"
SOFTWARE_VERSION_ID = "0.1"
INSTRUMENT_HOST_NAME = "MESSENGER"
INSTRUMENT NAME = "ENERGETIC PARTICLE AND PLASMA SPECTROMETER"
INSTRUMENT_ID = "EPPS"
DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
DATA SET NAME = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
MISSION PHASE NAME = "MERCURY ORBIT"
```

EPPSDDRSIS

TARGET NAME = "MERCURY"

START TIME = 2011-06-23T10:45:40.420458

STOP TIME = 2011-06-23T22:53:43.420605

SPACECRAFT CLOCK START COUNT = "1/217313408.800"

SPACECRAFT_CLOCK_STOP_COUNT = "1/217357091.000"

^HEADER = ("FIPS_FLUXMAP_2011174_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_FLUXMAP_2011174_DDR_V01.TAB", 4)

OBJECT = HEADER

 $\texttt{HEADER_TYPE} = \texttt{TEXT}$

INTERCHANGE_FORMAT = "ASCII"

RECORDS = 3

BYTES = 27486

DESCRIPTION = "The first four records of this

file are the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII TABLE

COLUMNS = 7

ROW BYTES = 9162

ROWS = 2

DESCRIPTION = "

This table contains angular flux map data as

described in EPPS DDR SIS."

^STRUCTURE = "FIPS_FLUXMAP_DDR.FMT"

```
END_OBJECT = ASCII_TABLE
```

END

5.3.4.9 FIPSKineticPropertiesLabel

AsampleFIPSKineticPropertiesDDRfilelabelisshownbelow:

MISSION_PHASE_NAME = "MERCURY ORBIT"

```
= "PDS3"
PDS VERSION ID
/* ** FILE FORMAT ** */
              = 42
FILE_RECORDS
          = FIXED_LENGTH
RECORD_TYPE
             = 176
RECORD BYTES
/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */
                     = "FIPS_NTP_2012054_DDR_V01"
PRODUCT_ID
PRODUCT_VERSION_ID = "01"
PRODUCT_CREATION_TIME = 2013-06-04T21:22:22
PRODUCT TYPE
                     = "DDR"
STANDARD_DATA_PRODUCT_ID = "FIPS_NTP_DDR"
                    = "mfips_ddr_sample.pro"
SOFTWARE NAME
SOFTWARE_VERSION_ID = "0.1"
INSTRUMENT_HOST_NAME = "MESSENGER"
INSTRUMENT_NAME = "FAST IMAGING PLASMA SPECTROMETER"
= "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"
DATA SET ID
DATA_SET_NAME
                      = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR
V2.0"
SOURCE PRODUCT ID = "FileByFile"
```

TARGET NAME = "MERCURY"

SPACECRAFT CLOCK START COUNT = "1/238523075.000"

START TIME = 2012-02-23T22:20:08.845

SPACECRAFT_CLOCK_STOP_COUNT = "1/238524872.000"

STOP_TIME = 2012-02-23T22:50:05.845

^HEADER = ("FIPS_NTP_2012054_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_NTP_2012054_DDR_V01.TAB", 4)

OBJECT = HEADER

HEADER TYPE = TEXT

INTERCHANGE_FORMAT = "ASCII"

RECORDS = 3

BYTES = 176

DESCRIPTION = "

This table contains FIPS ion number densities and temperatures,

as well as the pressure calculated from their product.

Quantities are calculated after sufficient time accumulation

to allow meaningful interpretation."

END OBJECT = HEADER

OBJECT = ASCII TABLE

COLUMNS = 13

ROW_BYTES = 176

ROWS = 39

DESCRIPTION = '

This table contains FIPS NTP values."

```
^STRUCTURE = "FIPS_NTP_DDR.FMT"
```

END_OBJECT = ASCII_TABLE

END

5.3.4.10 FIPSViewingNormalizationProductsLabels

5.3.4.10.1 FIPS Cartesian to MSO Coordinates Rotation Matrix (ROTMSO)

PDS_VERSION_ID = "PDS3"

/* ** FILE FORMAT ** */

FILE_RECORDS = 1303

RECORD_TYPE = FIXED_LENGTH

RECORD BYTES = 195

/* ** GENERAL DATA DESCRIPTION PARAMETERS ** */

PRODUCT_ID = "FIPS_ROTMSO_2010001_DDR_V01"

PRODUCT_VERSION_ID = "01"

PRODUCT_CREATION_TIME = 2014-06-04T10:00:00

PRODUCT_TYPE = "DDR"

STANDARD_DATA_PRODUCT_ID = "FIPS_ROTMSO_DDR"

SOFTWARE NAME = "MFIPS DDR SAMPLE.PRO"

SOFTWARE VERSION ID = "0.1"

INSTRUMENT HOST NAME = "MESSENGER"

INSTRUMENT NAME = "ENERGETIC PARTICLE AND PLASMA SPECTROMETER"

INSTRUMENT ID = "EPPS"

DATA_SET_ID = "MESS-E/V/H/SW-EPPS-3-FIPS-DDR-V2.0"

DATA_SET_NAME = "MESSENGER E/V/H/SW EPPS CALIBRATED FIPS DDR

V2.0"

MISSION PHASE NAME = "MERCURY 4 CRUISE"

TARGET NAME = "MERCURY"

SPACECRAFT_CLOCK_START_COUNT = "1/170791497.000"

START TIME = 2010-01-01T00:00:23.676

SPACECRAFT_CLOCK_STOP_COUNT = "1/170877732.000"

STOP_TIME = 2010-01-01T23:57:38.677

^HEADER = ("FIPS_ROTMSO_2010001_DDR_V01.TAB", 1)

^ASCII_TABLE = ("FIPS_ROTMSO_2010001_DDR_V01.TAB", 4)

OBJECT = HEADER

 $\texttt{HEADER_TYPE} = \texttt{TEXT}$

INTERCHANGE_FORMAT = "ASCII"

RECORDS = 3

BYTES = 585

DESCRIPTION = "The first three records of this

file are the header section. The header contains column

headings to improve usability."

END OBJECT = HEADER

OBJECT = ASCII TABLE

COLUMNS = 5

ROW BYTES = 195

ROWS = 1300

DESCRIPTION = "

This table contains rotation matrix from FIPS cartesian

to MSO for each energy scan."

^STRUCTURE = "FIPS_ROTMSO_DDR.FMT"

END OBJECT = ASCII TABLE

END

5.4 Standards Used in Generating Data Products

5.4.1 PDS Standards

The EPPSDDR data products are constructed according to the data object concepts developed by the PDS . By adopting the PDS format, the data products are consistent in content and organization with other planetary data collections . In the PDS standard, the DDR data file is grouped in to objects with PDS labels describing the objects. Each DDR data product consists of two files:

- AdatafilecontaininganASCIItableobject(theprimarydata),infixedfieldformat.ASCII
 tableobjectsareineithercommaseparatedvalue(CSV)format(EPS)orarewhitespace
 delimited(FIPS).ThismakestheASCIIdataextremelyeasytoreadbymanycommercialoffthe-shelfprograms.TheoneexceptionistheEPSPitchAngleSpectrogram,whichisabinary
 PNGfile.
- Alabelfilewhichservesasahigh-leveldescriptionoftheparametersofwhichcorrespond tothedatafile. The labelfile contains apointer to an external format file, which details the structure of the table object in the data file.

5.4.2 Time Standards

One of the time fields in the FIPS table objects references the Mission Elapsed Time (MET). This MET is the spacecraft time in integer seconds that is transmitted to MESSENGER subsystems by the Integrated Electronics Module (IEM). MET=0 is on August 3,2004, at 05:59:16 UTC (coordinated universal time), which is 1000 seconds prior to the MESSENGER launch. Relativistic effects and circumstances occurring during the mission would result in MET not being a true account of seconds since launch. Following a planned spacecraft clock reset (in early 2013, partition numbers (1/, or 2/) were added to product labels to disambiguate MET seconds after the spacecraft clock reset (if partition number is not present, SPICE defaults to partition 1/). For these reasons the MESSENGER spacecraft clock coefficients file is archived at the PDSNAIFN ode. This file is used in conjunction with the leap seconds kernel file in order to calculate the conversion between MET and UTC.

The conversion is easily done through the use of SPICE kernels and the CHRONOS Utility. CHRONOS is a utility included with the SPICE package that is distributed by the PDSNAIF node. The SPICE kernels are files that contain the information needed to perform the conversion. Two SPICE kernels are required. One is the Leap Seconds Kernel (LSK) and the other is the MESSENGERS pacecraft to the conversion of t

ClockKernel(SCLK). The SCLK file is used by CHRONOS to convert between spacecraft clock time and ephemeristime, while the LSK file is used to convert from ephemeristime to UTC time. The CHRONOS utility is self-documenting and the SPICE package itself contains full documentation on each of the utilities (including CHRONOS) and how they are used.

EPPSDDR data is time-tagged with spacecraft event time (SCET) in the following UTC format: CCYY-DDDTHH: MM: SS.sss. This format represents a concatenation of the conventional date and time expressions with the two parts separated by the letter T:

```
CC -century(00- 99)

YY -year(00- 99)

DDD -dayofyear(001-366)

T -date/timeseparator

HH-hour(00- 23)

MM -minute(00-59)

SS -second(00- 59)

sss-fractionsofsecond(000- 999)
```

5.4.3 Coordinate Systems

 $There are two coordinatesystems in use in the EPPSDDR data products: 1) the Mercury-centric Solar Orbital (MSO, defined in the MESSENGER SPICED ynamic Frames Kernel) used for spacecraft position vectors; and 2) the FIPSS pherical coordinatesystem, used for FIPS incident angless inceit represents natural coordinates for the sensor . The latter is a spherical version of the FIPS Cartesian coordinatesystem (FIPS_CART), which is defined in the MESSENGER SPICE Frames Kernel . FIPS Spherical coordinates (FIPS_SPH) consist of a radius (r), zenith angle (the ta) and a zimuth a langle (phi). The zenith angle is defined as the angle between the vector and the zaxis in the FIPS Cartesian coordinate system . It ranges from 0 to 180 degrees. The a zimuth a langle ranges from 0 to 360 degrees and is defined as the angle between the vector and the xaxis in the FIPS Cartesian coordinate system . The radius is defined as usual as the magnitude of the FIPS Cartesian vector.$

5.4.4 Data Storage Conventions

ThedataareorganizedfollowingPDSstandardsandstoredonharddiskattheMESSENGERSOC.
TheSOCtransfersdatatoPDSasdetailedinsection 5.3.3. AfterverificationofthedatatransferPDS

provides public access to MESSENGER science data products through its online data distribution system.

5.5 Data Validation

The EPPSDDR data archive volume set includes all data acquired during the MESSENGER mission. The archive validation procedure described in this section applies to data products generated during all post launch phases of the mission. To be clear, there is one and only one documentation volume and on ly one EPPS data archive volume created over the whole mission. Release dates are stated in the schedule in section. The data volume occurred according to the same schedule. Updates to the documentation volume occurred according to this schedule and at the discretion of the EPPS team.

PDS standards recommend that all data included in the formal archive bevalidated through a peer-review process. This process is designed to ensure that both the data and documentation are of sufficient quality to be useful to future generations of scient ists. The process is presented as several steps, most of which occur in the PDS peer review. This peer review is conducted before any volumes are produced and released to PDS.

The peer review panel consists of members of the EPPS team, the PPI node of PDS, and at least one outsides cientist actively working in the field of energetic particles research. The PDS personnel are responsible for validating that the volumes are fully compliant with PDS standards. The instrument team and outside reviewer (s) are responsible for verifying the content of the dataset, the completeness of the documentation, and the usability of the data in its archive format.

The peer review validates the document at ion and data archive volumes. First the panel review sthis document and verifies that the volumes and DDRs produced to this specification will be useful. The peer review also validates the EPPSDDR data in a two step process. The first step consists of reviewing a sample data set for compliance with the PDS standards. The sample data set is delivered and reviewed in conjunction with delivery and review of this SIS document. The second step is examination of the data to ensure usability and completeness. The PDS personnel are responsible for validating that the DDR data set is fully compliant with PDS standards. The instrument team and the outsides cience reviewer (s) are responsible for verifying the content of the data set, the completeness of the document at ion, and the usability of the data in its archive format.

Any deficiencies in the archive data or documentation volumes are recorded as liens against the action of the contraction ofproductbythereviewpanel. The sampled at a set is created using software provided by APL and the University of Michigan. Once the sample data are validated, and all liens placed against the product orproductgenerations of tware are resolved, the sames of tware will be used to generate subsequentdataproductsinanautomatedfashion.

Duringautomated production, the data file content is spotchecked by members of the EPPS team."Quick look" products generated by software provided by ACT and the EPPS team areproduced routinelyandexaminedbymembersoftheteam.Inaddition,thedataareactivelyusedbyteam memberstoperformtheiranalysis. Any discrepancies in the data noted during these activities will be investigated. If the discrepancy is a data error, the response depends on the source of the error.If the error is in the software producing the data product, the error iscorrectedandthedata affected isreproduced, replacing the data file. If there is a correct able error in a data file, the file is replaced. If an error in a data file is uncorrectable, the error isdescribedinthecumulativeerrata fileincludedinthearchivevolume. The structure of data files and labels will be spotchecked by the PPInodeforcompliancewithPDSstandardsandthisSIS.

Detailed Data Product Specification

6.1 Data Product Structure and Organization

The MESSENGERE PPSDDR data products are archived at the PDSPPIN ode. The automated account of the product ofproduction and release of DDRs lends itself to the regular release schedule outlined in section7.If errors are discovered the data are replaced with corrected DDRs on the next scheduled deliverydate.

Calibrationtables and calibration procedures are required to properly analyze DDRs. These ancillary data are archived at the PDSPPIN ode as part of the EPPS document at ion volume. The ancillary data are archived at the PDSPPIN ode as part of the EPPS document at ion volume. The property of the PDSPPIN ode as part of the EPPS document at ion volume. The property of the EPPS document at ion volume are also as a part of the EPPS document at ion volume. The property of the EPPS document at ion volume at ion volume and the EPPS document at ion volume. The experiment at ion volume are also as a part of the EPPS document at ion volume at ion volume. The experiment at ion volume at ion voludocumentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes are referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes are referenced by all EPPS data archive volumes. The documentation volume is referenced by all EPPS data archive volumes are referenced by a $therefore includes the {\tt EPPSEDRSIS}, the {\tt EPPSCDRSIS}, and the {\tt EPPSDDRSIS} in addition to the {\tt EPPSCDRSIS}.$ calibrationtables, calibration procedures, and other documents applicable to the data archive volumes. A first release of the EPPS documentation volume accompanied the initial release of the accompanied the accompEPPSEDRdataarchive.AnupdatetotheEPPSdocumentationvolumewillaccompanytheinitial release of the DDR data archive. After the initial releases of the DDR level documentation there wereupdatestothedocumentationvolumetodocumentdataqualityandasneededforproduct and calibration updates.

6.2 Handling Errors

The possibility exists that errors may be introduced into the archive even with validation procedures applied to the archive volumes. An ERRATA report file is maintained to track and document all discovered uncorrectable errors that may occur during the mission. Correctable errors, such as revised DDRs or DDRs that were missing from a previous PDS delivery are provided at the next scheduled PDS delivery or at the final delivery date (schedule in section 7). PDS then replaces the outdated files with the revised DDR files in the data directories of the archive volume. Filer evisions are also recorded in the data product labelity words PRODUCT_VERSION_ID and PRODUCT_CREATION_TIME, which can be used in addition to ERRATA. TXT to detect updates. The ERRATA report file is archived in the ROOT directory of the EPPSDDR data volume.

6.3 Data Format Description

Data are stored in ASCII table format. A detached PDS label file provides a detailed description of the structure of each table.

6.4 Label and Header Descriptions

The following are the keyword definitions for the detached PDS label file, which accompanies the instrument data file. The detached PDS label file has the same name as the data file it describes, except for the extension. LBL to distinguish it as a label file.

PDS VERSION ID

Represents the version number of the PDS standards documents that is valid when a data product label is created. PDS 3 is used for the MESSENGER data products.

FILE_RECORDS

Indicates the number of physical file records, including both label records and data records.

RECORD TYPE

 $Indicates the record format of a file \qquad . \ Note: In the PDS, when record_type is used in a detached label file it always describes its corresponding detached data file, not the$

 $label file itself\ .\ The use of record_type along with other file-related data elements is fully described in the PDSS tandards Reference.$

RECORD_BYTES

 $Indicates the number of bytes in a physical file record, including record terminators and separators. Note: In the PDS, the use of record_bytes, along with other file-related data elements is fully described in the Standards Reference.\\$

PRODUCT_ID

Represents a permanent, unique identifier as signed to a data product by its producer.

PRODUCT_CREATION_TIME

Defines the UTC system format time when a product was created.

PRODUCT_VERSION_ID

Identifies the version of an individual product within a dataset.

Example: 1.0, 2.0, 3.0 .

 $Product_version_idisincremente difagiven DDR has to be regenerated and sent to PDS to replace a previously submitted DDR.$

PRODUCT TYPE

Identifies the type or category of a product within a dataset.

STANDARD_DATA_PRODUCT_ID

Used to link an EPPSDDR file to one of the 12 types of EPPS data products defined within the EPPSDDRS IS.

SOFTWARE_NAME

Identifies the data processing software used to convert from CDR into DDR products.

SOFTWARE_VERSION_ID

Indicates the version of the data processing software used to generate the DDR products.

MD5_CHECKSUM

Used to verify the successful electronic transfer of the DDR from the SOC to the PDS-PPIN ode.

INSTRUMENT_HOST_NAME

 $The full name of the host on which an instrument is based \\ MESSENGER spacecraft.$. In this case it is the MESSENGER spacecraft.

INSTRUMENT_NAME

Providesthefullnameoftheinstrument.

INSTRUMENT_ID

Provides an abbreviated name or a cronymwhich identifies an instrument.

DATA_SET_ID

Thedata_set_idelementisauniquealphanumericidentifierforadatasetoradata product.Thedata_set_idvalueforagivendatasetorproductisconstructedaccording toflightprojectnamingconventions . Thereisonlyonedata_set_idfortheEPPSDDRs.

MISSION PHASE NAME

Provides the commonly used identifier of a mission phase.

TARGET_NAME

Thetarget_nameelementidentifiesatarget . Thetargetmaybeaplanet, satellite, ring, region, feature, asteroidor comet.

START_TIME

Provides the date and time of the beginning of an event or observation (whether it be a spacecraft, ground-based, or system event) in UTCs ystem form at .

STOP_TIME

 $Provides the date and time of the end of an observation or event (whether it be a spacecraft, ground-based, or system event) in UTCs ystem format \\ .$

SPACECRAFT_CLOCK_START_COUNT

Providesthevalueofthespacecraftclockatthebeginningofatimeperiodofinterest

SPACECRAFT_CLOCK_STOP_COUNT

Provides the value of the space craft clock at the end of a time period of interest

^TABLE

PointertotheDDRfile, which contains the data in ASCII table format. The structure of the data file is defined in a reference of the data file.

OBJECT

 $Specifies that the DDR is a PDSTABLE object. This object contains its own elements, which are defined below. NOTE: the end of the object definition is always marked with an END_OBJECT line.\\$

COLUMNS

Identifies the number of columns (fields) in the table.

INTERCHANGE_FORMAT

This element specifies that the table is in ASCII format.

ROW_BYTES

Specifies the number of bytes for each row in the table.

ROWS

Identifies the number of rows (records) in the table.

^STRUCTURE

This is a pointer to the external file which provides the structure definition for the table object.

The following describes the keywords used to describe the PDST able Object. These keywords are contained in the FORMAT (.FMT) files for each DDR data product.

COLUMN_NUMBER

 $Identifies the location of the column within the larger data object (such a satable). For tables consisting of rows (I=1,N) and columns (j=1,M) the column_number is the j-th index of any row. \\$

NAME

Indicates a literal value representing the common term used to identify an element or object. NOTE: in the PDS data dictionary, name is restricted to 30 characters and must conform to PDS no menclature standards.

BYTES

Specifies the number of bytes allocated for this particular column element.

DATA TYPE

Specifies the internal representation and/or mathematical properties of the value being stored in this column.

START_BYTE

Identifies the location of the first byte of the particular column, counting from 1.

ITEMS

Defines the number of multiple, identical occurrences of a single object. Used mainly in columns containing spectral or histogram data.

ITEM_BYTES

 $The size in bytes of individual items in a column. ITEMS*ITEM_BYTES should equal the value in the BYTES column.\\$

The format file contains the full text for describing each column of the table. See Appendices for listing of each field in the individual format files.

6.5 File Naming Conventions

The filenames developed for PDS data volumes are restricted to a maximum 36-character filename and 3-character extension name with a period separating the file and extension names.

 $The general form of the EPPSEPS file name for pitch angle data is \\ \text{``EPSP_AyyyydddDDR_V\#.TAB''} \\ where$

EPS Instrumentname

P_A PitchAngle

yyyy Fourdigityear

ddd Threedigitdayofyear

DDR CODMACprocessinglevel

V# Versionnumber

 $The general form of the {\tt EPPSEPS} file name for pitch angle spectrograms is$

"EPS_PAS_yyyydddhhmmss _V#.PNG" where

EPS Instrumentname

PAS PitchAngleSpectrogram

a

yyyy Fourdigityear

ddd Threedigitdayofyear

hhmmss Hour, minute, second

V# Versionnumber

PNG PortableNetworkGraphicsfileextension

The date in the filename is the start time of the data contained in the file.

ThegeneralformoftheEPPSFIPSfilenameforDDRs is "FIPS_<TYPE>_yyyyddd_DDR_V#.TAB" where

FIPS Instrumentname

 $\verb| <TYPE> Refers to the type of data contained in the file. Possible values are$

ESPEC-EnergySpectra

NOBS -ObservedDensity

PCHANG -PitchAngle

ERPCHANG - Energy-Resolved Pitch Angle

ARRDIR -ArrivalDirection

FLUXMAP - Angular Flux Map

NTP -KineticProperties

ROTMSO -ViewingNormalizationrotationmatrixtoMSO

yyyy Fourdigityear

ddd Threedigitdayofyear

DDR CODMACprocessinglevel

V# Versionnumber

ForallEPPSdata, the initial version number is "V1" (EPS and FIPS_FOVPIX) or "V01" (all other FIPS). The version number increments for each successive version of the DDR product that is produced. A new version of the DDR product may be produced as a result of an error in the product or a sare sult of errors discovered in the product generation process.

For all EPPS data except the EPS Pitch Angle Spectrogram:

TAB the file extension is dependent on the file type

.TAB, EPS and FIPS Instrument Data in ASCII table

.LBL, Detached PDS label file

6.6 Archive Volume and File Size

Twoarchivevolumesarecreatedtoarchiveboththe EPPSDDR data and the documentation, which is needed to analyze the DDRs. The first volume is the EPPSD ocumentation Volume, having volume IDMESSEPPS_DOC. This documentation volume contains products related to the EPPSEDR, CDR and DDR data archives including:

- 1. AllrequiredPDScatalogfilesfortheEDR,CDR,andDDRarchives.
- 2. TheEDR,CDR,andDDRSISdocuments.
- $3. \quad The Space Sciences Review (SSR) in strument paper once copyright permission is obtained. \\ This may not be included in the initial release for copyright reasons.$
- 4. The EPPS calibration report.
- 5. The EPPS calibration procedures document.
- 6. Calibrationtables.
- $7. \quad Other documents considered useful by the {\tt MESSENGER} project or the {\tt EPPS} team.$

 $The data archive volume, designated the EPPSD at a Archive Volume and having volume ID MESSEPPS_DDR, contains the DDR data and required files for conforming to PDS volume archive standards. This includes the index files, AAREADME. TXT file, etc. The approximate data archive volume size is 11 GB. \\$

6.7 Directory Structure and Contents for EPPS Documentation Volume

 $The following illustrations how sthe directory structure overview for the {\tt EPPS} documentation volume.$

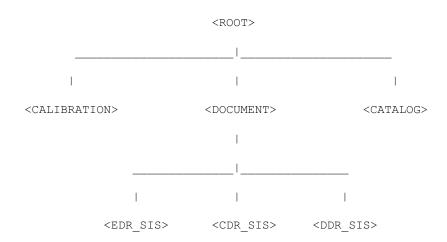


Figure 2 Documentation Volume Structure

6.7.1 Directory Contents

<ROOT>Directory

This is the top-level volume directory. The following are files contained in the root directory.

AAREADME.TXT - General information file. Provides users with an overview of the contents and organization of the associated volume, general instructions for its use, and contact information.

VOLDESC.CAT - PDS file containing the VOLUME object. This gives a high-level description of the contents of the volume. Information includes: production date, producer name and institution, volume ID, etc.

ERRATA.TXT - Text file for identifying and describing errors and/or anomalies found in the current volume, and possibly previous volumes of a set. Any known errors for the associated volume are documented in this file.

<CALIBRATION>Directory

This contains the calibration tables needed to analyze the EPPS CDR data. The calibration tables are in ASCII format. Format files for the calibration tables are also located here, as are the following files.

CALINFO.TXT-Brief description of the directory contents and naming conventions.

EPPS_*_EDR2CDR.PDF: Describes the procedure used to convert EDRs to CDRs for each instrument, (as indicated by the * text).

FIP*.TAB:TheFIPSenergyperchargetables.

<CATALOG>Directory

This subdirectory contains the catalog object files for the entire volume. The following files are included in the catalog subdirectory.

CATINFO.TXT: Identifies and describes the function of each file in the catalog directory.

EPPS*DATASET.CAT: Describes the general content of the EDR data set for each instrument (as indicated by the * text) and includes information about the duration of the mission and the person or group responsible for producing the data.

EPPS*DATASET_CDR.CAT: Describes the general content of the CDR data set for each instrument (as indicated by the * text) and includes information about the duration of the mission and the person or group responsible for producing the data.

EPPS*DATASET_DDR.CAT: Describes the general content of the DDR data set for each instrument (as indicated by the * text) and includes information about the duration of the mission and the person or group responsible for producing the data.

INSTRUMENT.CAT: Describes physical attributes of the EPPS instrument and provides relevant references to published literature.

INSTHOST.CAT: Describes the MESSENGER spacecraft.

MISSION.CAT: Describes the scientific goals and objectives of the MESSENGER program. It also identifies key people and institutions.

PERSON.CAT: Lists and provides contact information for the people involved in the MESSENGER mission, including those involved with EPPS.

REF.CAT: Provides references to scientific papers and other publications of interest to those using the data, both for EPPS and the mission as a whole.

< DOCUMENT > Directory

This subdirectory contains the documentation that is needed in order to understand and analyze the EDR, CDR, and DDR data volumes. The documents are separated into individual subdirectories according to the document type. The document types are not restricted to the four shown in the graphical depiction of the directory structure. There are additional document types as needed to categorize each document. The following file is included in the subdirectory.

DOCINFO.TXT: Identifies and describes the function of each file in the DOCUMENT directory.

6.8 Directory Structure and Contents for EPPS Data Volume

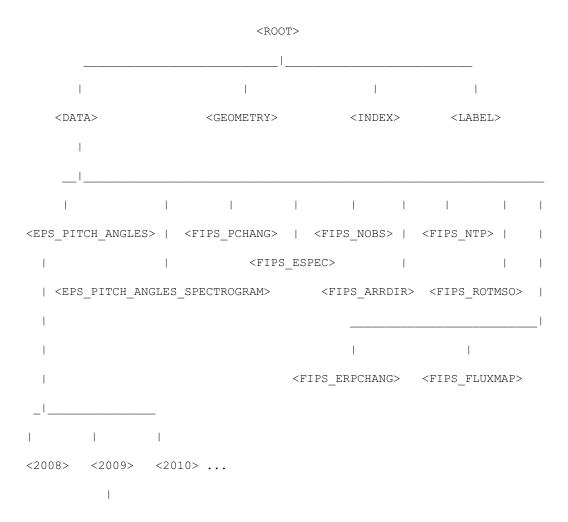




Figure 3 Data Volume Directory Structure

6.8.1 Directory Contents

<ROOT>Directory

This is the top-level directory of a volume. The following are files contained in the root directory.

AAREADME.TXT - General information file. Provides users with an overview of the contents and organization of the associated volume, general instructions for its use, and contact information.

VOLDESC.CAT - PDS file containing the VOLUME object. This gives a high-level description of the contents of the volume. Information includes: production date, producer name and institution, volume ID, etc.

ERRATA.TXT - Text file for identifying and describing errors and/or anomalies found in the current volume, and possibly previous volumes of a set. Any known errors for the associated volume are documented in this file. This includes revised DDRs meant to replace DDRs in a previous PDS delivery.

<DATA>Directory

 $This top level directory contains the subdirectories corresponding to these vendata products (section 5.2) and supporting products. The directories are further subdivided into YEAR and MONTH directories. The FIPS_FOVPIX directory contains the Pixel Field Of View ancillary products. \\$

<GEOMETRY>Directory

This subdirectory contains information about the files (e.g. SPICE kernels, etc) needed to describe the observation geometry for the data.

GEOMINFO.TXT:Identifies and describes the SPICE kernels that a user must have in order to determine observation geometry for the data. The SPICE kernel files are archived with the PDSNAIF node.

<INDEX>Directory

This subdirectory contains the indices for all data products on the volume. The following files are contained in the index subdirectory.

INDXINFO.TXT – Identifies and describes the function of each file in the index subdirectory. This includes a description of the structure and contents of each index table in the subdirectory AND usage notes.

INDEX.TAB - The DDR index file is organized as a table: there is one entry for each of the data files included in the EPPS data set; the columns contain parameters that describe the observation and instrument and spacecraft parameters. These parameters include state information, such as integration time, spacecraft clock count, time of observation, and instrument modes.

INDEX.LBL - Detached PDS label for INDEX.TAB. It contains the INDEX_TABLE object which identifies and describes the columns of the EPPS index table.

MD5.TAB - The MD5 checksum file that contains MD5 hash values for every file in the volume.

MD5.LBL - Detached PDS label for MD5.TAB.

<LABEL>Directory

This subdirectory contains the "label fragments" (i.e., the *.FMT files) for all data products on the volume. These format files describe the table and data objects which store the data.

7 Archive Release Schedule to PDS

TheMESSENGEREPPSdataandvolumearchivesweretransferredfromtheSOCtothePDSPPI Nodeusingtheprocessdetailedinsection 5.3.3.TheSPICEkernelsweretransferredtotheNAIF node.Thetransfertookplaceaccordingtotheschedulein[2].

8 Appendices

8.1 EPS PITCH ANGLES.FMT Table Columns

 $The following are the columns as defined by the EPS_PITCH_ANGLES.FMT structure file. This file defines the ASCII table containing the EPS pitch angled at a. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a particular column in the table at a type for a type for a particular column in the table at a type for a particular column in the table at a type for a$

Table 3 EPS_ PITCH_ANGLES.FMT Columns

Length (bytes)	Data Type	Column Name	Summary (see full text for column description)
21	TIME	TIME	Spacecrafteventtime(UTC)for this data record.
22	ASCII_REAL	PITCH_ANGLE_S0	Pitchangle (degrees) forsector 0.
22	ASCII_ REAL	PITCH_ANGLE_S1	Pitchangle(degrees)forsector 1.
22	ASCII_ REAL	PITCH_ANGLE_S2	Pitchangle(degrees)forsector 2.
22	ASCII_ REAL	PITCH_ANGLE_S3	Pitchangle(degrees)forsector 3.
22	ASCII_ REAL	PITCH_ANGLE_S4	Pitchangle(degrees)forsector 4.
22	ASCII_ REAL	PITCH_ANGLE_S5	Pitchangle(degrees)forsector 5.

8.2 FIPS_PCHANG_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS \\ - PCHANG_DDR. FMT structure file. This file defines the ASCII table containing the FIPS pitch angled at a. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table.$

 $The INDEX field values are unique for a given day and may be used to match data in this file to that in the other FIPSDDR files for the same day. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only in the FIPS_NOBS_DDR file.$

Table 4 FIPS_PCHANG_DDR.FMT

Length (bytes)	Data Type	Field Name	Summary (see full text for column description)
7	ASCII Integer	INDEX	Uniqueidentifierforthecurrentdatasample.
14	ASCIIReal	MET	Mission Elapsed Time in seconds.
554	ASCIIReal	H_PA	H+flux -pitchanglehistogram.Pitchanglerange=0 - 180inclusive,where0isparalleltomagneticfield. Binsizeis5degrees .
554	ASCIIReal	HE2_PA	He2+flux -pitchanglehistogram.Pitchanglerange= 0 - 180inclusive,where0isparalleltomagnetic field.Binsizeis5degrees.
554	ASCIIReal	HE_PA	He+flux -pitchanglehistogram.Pitchanglerange= 0 - 180inclusive,where0isparalleltomagnet ic field.Binsizeis5degrees.
554	ASCIIReal	NAGROUP_PA	Na+groupflux -pitchanglehistogram.Pitchangle range=0 - 180inclusive,where0isparallelto magneticfield.Binsizeis5degrees.
554	ASCIIReal	OGROUP_PA	O+groupflux -pitchanglehistogram.Pitchangle range=0 - 180inclusive,where0isparallelto magneticfield.Binsizeis5degrees.

8.3 FIPS_ERPCHANG_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_ERPCHANG_DDR. FMT structure file. This file defines the ASCII table containing the FIPS energy-resolved pitch angled at a. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

 $The START_INDEX and STOP_INDEX field values are unique for a given day and may be used to match data in this file to that in the other FIPSDDR files for the same day. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only in the FIPS_NOBS_DDR file.$

Table 5 FIPS_ERPCHANG_DDR.FMT

Length (bytes)	Data Type	Field Name	Summary (see full text for column description)
14	ASCIIInteger	START_INDEX	Uniqueidentifierforthe startofthecurrent samplerange.
14	ASCIIInteger	STOP_INDEX	Uniqueidentifierforthe endofthecurrent samplerange.
14	ASCIIReal	START_MET	$\label{lem:missionElapsedTimeinseconds} \mbox{ at the beginning } \\ \mbox{ of the accumulation }.$
14	ASCIIReal	STOP_MET	MissionElapsedTimeinseconds attheendofthe accumulation.
16	CHARACTER	TIME_RESL	Timeresolutionlabel.
16	CHARACTER	ION	Iongrouplabel.
1152	ASCIIReal	ERPCHANG	2Dmatrix(Pitchangle,DSHVstep).Pitchangle range:0 - 180inclusive dividedover10degree bins.Matrixisorderedbysequential64element pitchanglevectors(elements0 -63 arepitchangle 0,DSHVsteps0to63; elements64 -127arepitch angle10,steps0to63;andsoon).

8.4 FIPS_ESPEC_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_ESPEC_DDR. FMT structure file. This file defines the ASCII table containing the FIPS energy spectra data. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table e. \\$

The INDEX field values are unique for a given day and may be used to match data in this file to that in the other two FIPSDDR files for the same day and the first of the same day and the first of the same day and the first of the first o

Table 6 FIPS_ESPEC_DDR.FMT Columns

Length	Data Type	Field Name	Summary (see full text for column
(bytes)			description)
7	ASCII Integer	INDEX	Uniqueidentifierforthecurrentdatasample.
14	ASCII Real	MET	Mission Elapsed Time in seconds.
959	ASCIIReal	Н	H+ fluxpere/q in1/(cm 2 skV).
959	ASCIIReal	HE2	He2+fluxpere/q in1/(cm 2 skV).
959	ASCIIReal	НЕ	He+ fluxpere/q in1/(cm 2 skV).
959	ASCIIReal	NA_GROUP	Na+ group fluxpere/q in1/(cm^2 skV).
959	ASCIIReal	O_GROUP	O+ group fluxpere/q in1/(cm 2 skV).

8.5 FIPS NOBS DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_NOBS_DDR. FMT structure file. This file defines the ASCII table containing the FIPS observed density data. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

The INDEX field values are unique for a given day and may be used to match data in this file to that in the other two FIPSDDR files for the same day and the same day. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only this file.

Table 7 FIPS_NOBS_DDR.FMT Columns

Length	Data Type	Field Name	Summary (see full text for column
(bytes)			description)
7	ASCII Integer	INDEX	Uniqueidentifierforthecurrentdata sample.
14	ASCIIReal	MET	Mission Elapsed Time in seconds.
7	ASCIIReal	ACCUM	$\label{lem:complete} Accumulation time for the current data sample.$
15	ASCIIReal	YFR	Timeatendoftheaccumulationinfloatingpoint year.
8	ASCII Real	DOYFR	Timeatendoftheaccumulationin floatingpoint dayofyear .
5	ASCII Integer	HOURS	Hourattheendoftheaccumulation .
7	ASCII Integer	MINUTES	Minuteattheendoftheaccumulation .
7	ASCII Real	SECONDS	Secondattheendoftheaccumulation .
10	ASCIIReal	MSOX	MESSENGERX -coordinate in MSGR_MSO frame in km.
10	ASCIIReal	MSOY	MESSENGERY -coordinateinMSGR_MSOframe inkm.
10	ASCIIReal	MSOZ	$\label{lem:messenger} MESSENGERZ\ -coordinate in MSGR_MSO frame in km.$
6	ASCII Real	LAT	MESSENGERpositioninMercurylatitudein degrees.
6	ASCII Real	MLT	$\label{lem:messenger} MESSENGER position in magnetic local time in hour fraction\ .$
9	ASCII Real	ALT	MESSENGERaltitudeinkm .
14	ASCIIReal	Н	H+ observeddensityincm -3.
14	ASCIIReal	HE2	He2+ observeddensityincm -3.
14	ASCIIReal	НЕ	He+ observeddensityincm -3.

14	ASCIIReal	NA	Na+ observeddensityincm -3.
14	ASCIIReal	0	0+ observeddensityincm -3.
4	ASCII Integer	QUAL	Dataqualityflag(0=good,1=bad) .

8.6 FIPS_ARRDIR_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_ARRDIR_DDR.FMT structure file. This file defines the ASCII table containing the FIPS arrival direction data. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

 $The START_INDEX and STOP_INDEX field values are unique for a given day and may be used to match data in this file to that in the other FIPSDDR files for the same day. For example, the START_INDEX corresponds to the INDEX for the first line in the FIPSNOBS data used for this accumulation. The STOP_INDEX corresponds to the last line. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only this file.$

Table 8 FIPS_ARRDIR_DDR.FMT Columns

Length	Data	Field Name	Summary (see full text for column
(bytes)	Туре		description)
14	ASCII Integer	START_INDEX	Unique identifierforthestartofthecurrent samplerange .
14	ASCII Integer	STOP_INDEX	Uniqueidentifierfortheendofthecurrent samplerange.
14	ASCII Real	START_MET	$\label{lem:missionElapsedTimeat} Mission Elapsed Time at the beginning of the accumulation.$
14	ASCII Real	STOP_MET	MissionElapsedTimeattheendofthe accumulation.
14	ASCII	TIME_RESL	String label for time resolution of current line.
14	ASCII	ION	Nameofionoriongroup .

9072	ASCII	DIRECTIONAL_FLUX	Ionfluxasafunctionofarrivaldirection.Units:
	Real		(cm^2skeV/keVsr)^ -1.

8.7 FIPS_FLUXMAP_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_FLUXMAP_DDR. FMT structure file. This file defines the ASCII table containing the FIPS arrival direction data. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

 $The START_INDEX and STOP_INDEX field values are unique for a given day and may be used to match data in this file to that in the other FIPSDDR files for the same day. For example, the START_INDEX corresponds to the INDEX for the first line in the FIPSNOBS data used for this accumulation. The STOP_INDEX corresponds to the last line. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only this file. \\$

Table 9 FIPS_FLUXMAP_DDR.FMT Columns

Length	Data	Field Name	Summary (see full text for column
(bytes)	Type		description)
14	ASCII Integer	START_INDEX	Uniqueidentifierforthestartofthecurrent samplerange.
14	ASCII Integer	STOP_INDEX	Uniqueidentifierfortheendofthecurrent samplerange.
14	ASCII Real	START_MET	$\label{lem:missionElapsedTimeatthebeginning} Mission Elapsed Time at the beginning of the accumulation.$
14	ASCII Real	STOP_MET	MissionElapsedTimeattheendofthe accumulation.
14	ASCII	TIME_RESL	String label for time resolution of current line.
14	ASCII	ION	Nameofionoriongroup.
9072	ASCII Real	DIRECTIONAL_FLUX	Ionfluxasafunctionofarrivaldirection.Units: (cm^2skeV/keVsr)^ -1.

8.8 FIPS_NTP_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_NTP_DDR. FMT structure file. This file defines the ASCII table containing the FIPS arrival direction data. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

 $The START_INDEX and STOP_INDEX field values are unique for a given day and may be used to match data in this file to that in the other FIPSDDR files for the same day. For example, the START_INDEX corresponds to the INDEX for the first line in the FIPSNOBS data used for this accumulation. The STOP_INDEX corresponds to the last line. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only this file.$

Table 10 FIPS_NTP_DDR.FMT Columns

Length	Data Type	Field Name	Summary (see full text for column description)
(bytes)			
14	ASCII Integer	START_INDEX	Uniqueidentifierforthestartofthecurrent samplerange.
14	ASCII Integer	STOP_INDEX	Uniqueidentifierfortheendofthecurrent samplerange.
14	ASCIIReal	START_MET	$\label{lem:missionElapsedTimeatthebeginning} Mission Elapsed Time at the beginning of the accumulation.$
14	ASCIIReal	STOP_MET	MissionElapsedTimeattheendofthe accumulation.
14	ASCII	TIME_RESL	Stringlabel for time resolution of current line.
14	ASCII	ION	Nameofionoriongroup.
14	ASCIIReal	N	Ionnumberdensityinunitsofcm^ -3.
14	ASCIIReal	T	IontemperatureinunitsofMK .
14	ASCIIReal	P	IonpressureinunitsofnPa.

14	ASCIIReal	N_ERR	Errorinionnumberdensityinunitsofcm^ -3.
14	ASCIIReal	T_ERR	Error inion temperature in units of MK.
14	ASCIIReal	P_ERR	ErrorinionpressureinunitsofnPa
6	ASCII Integer	QUAL	QualityFlag .0=Good;non -0=Bad.Non -zeroflag valuesTBD.

8.9 FIPS_ROTMSO_DDR.FMT Table Columns

 $The following are the columns as defined by the FIPS_ROTMSO_DDR. FMT structure file. This file defines the ASCII table containing the FIPS Cartesian to MSO rotation matrix definition. The archive volume is optimized by defining the table structure once and providing are ference to it in the PDS label file. The columns are numbered according to their column order in the table. Data Type refers to the PDS standards data type for a particular column in the table. \\$

The INDEX field values are unique for a given day and may be used to match data in this file to that in the other two FIPSDDR files for the same day. This fact is used to avoid data duplication: Universal Time (UTC) and MESSENGER positions are given only this file.

Table 11 FIPS_ROTMSO_DDR.FMT Columns

Length	Data Type	Field Name	Summary (see full text for column
(bytes)			description)
11	ASCII Integer	INDEX	Uniqueidentifierforthecurrentdatasample.
20	ASCIIReal	MET	MissionElapsedTimeatthe end ofthe correspondingFIPSscan .
54	ASCIIReal	MATRIX_ROW_0	FIPSFOVtoMSGRMSO rotationmatrixrow0 .
54	ASCIIReal	MATRIX_ROW_1	$FIPSFOV to MSGRMSO rotation matrix row 1 \\ \\ .$
54	ASCIIReal	MATRIX_ROW_2	FIPSFOVtoMSGRMSOrotationmatrixrow2 .

8.10 SPICE Kernel Files Used in MESSENGER Data Products

The following SPICE kernel files were used to compute the UTC time and any geometric quantities found in the PDS labels . Kernel files were generated throughout the mission with a file naming convention specified by the MESSENGER project.
*.bsp:
$\label{lem:messenger} MESSENGER spacecrafte phemeris file. \ Also known as the Planetary Spacecraft Ephemeris Kernel (SPK) file.$
*.bc:
MESSENGER space craftorient at ion file. Also known as the Attitude C-Kernel (CK) file.
*.tf:
MESSENGER reference frame file. Also known as the Frames Kernel. Contains the MESSENGER spacecraft, science in strument, and communications antenna e frame definitions.
*.ti:
MESSENGER instrument kernel (I-kernel). Contains reference stomounting alignment, operation modes, and timing as well as internal and field of view geometry for the EPPS.
*.tsc:
MESSENGER spacecraft clock coefficients file. Also known as the Spacecraft Clock Kernel (SCLK) file.
*.tpc:

Planetary constants file. Also known as the Planetary Constants Kernel (PcK) file.

*.tls:

 $NAIF leap seconds kernel file. Used in conjunction with the SCLK kernel to convert between \\ Universal Time Coordinated (UTC) and MESSENGER Mission Elapsed Time (MET). Also called the \\ Leap Seconds Kernel (LSK) file.$

8.11 CODMAC/NASA Definition of Processing Levels

CODMAC/NASADefinition of processing levels for science datasets

CODMAC Level	Proc.Type	DataProcessingLevelDescription
1	RawData	Telemetrydatastreamasreceivedatthegroundstation,with scienceandengineeringdataembedded.CorrespondstoNASA packetdata.
2	Edited Data	Instrumentsciencedata(e.g.rawvoltages,counts)atfull resolution,timeordered,withduplicatesandtransmission errorsremoved.ReferredtointheMESSENGER programas Experiment DataRecords(EDRs).CorrespondstoNASALevel0 data.
3	Calibrated Data	Editeddatathatarestillinunitsproducedbyinstrument,but havetransformed(e.g.calibrated,rearranged)inareversible mannerandpackagedwithneededancillaryandauxiliarydata (e.g.radianceswithcalibrationequationsapplied).Referr edto intheMESSENGERProgramasCalibratedDataRecords(CDRs). Insomecasesthesealsoqualifyasderiveddataproducts (DDRs).CorrespondstoNASALevel1A.
4	Resampled data	Irreversiblytransformed(e.g.resampled,remapped,calibrated) valuesofth einstrumentmeasurements(e.g.radiances, magneticfieldstrength).ReferredtointheMESSENGER programaseitherderiveddataproducts(DDPs)orderived analysisproducts(DAPs).CorrespondstoNASALevel1B.
5	DerivedData	Derivedresultssuchasmaps ,reports,graphics,etc. CorrespondstoNASALevels2through5

CODMAC Level	Proc.Type	DataProcessingLevelDescription
6	Ancillary Data	Non-Sciencedataneededtogeneratecalibratedorresampled datasets.Consistsofinstrumentgains,offsets;pointing informationforscanplatforms,etc.
7	Corrective Data	Othersciencedataneededtointerpretspace -bornedatasets. Mayincludegroundbaseddataobservationssuchassoiltype oroceanbuoy measurementsof winddrift.
8	User Description	Descriptionofwhythedatawererequired,anypeculiarities associatedwiththedatasets,andenoughdocumentationto allowsecondaryusertoextractinformationfromthedata.

The above is based on the national research council committee on data management and computation (CODMAC) datalevels.

8.12 MESSENGER Glossary and Acronym List

ACT AppliedCoherentTechnologyCorporation AMU AtomicMassUnit APL The Johns Hopkins University Applied Physics LaboratoryASCII American Standard Code for Information InterchangeCDR CalibratedDataRecord CK AttitudeC-Kernel(SPICE) CODMAC Committee on Data Management and ComputationDAP DerivedAnalysisProducts DerivedDataProducts DDP **EDR** ExperimentDataRecords **EPPS** Energetic Particle and Plasma Spectrometer**EPS** EnergeticParticleSpectrometer

ESA ElectrostaticAnalyzer

FIFO FirstIn, Firstout. An electronic component that stores and retrieves

informationfollowingafirst- in-first-outdiscipline.

FIPS FastImagingPlasmaSpectrometer

FOV Field- of-View

FSW FlightSoftware

FTP FileTransferprotocol

GEANT4 GEometryANdTrackingsoftwaretoolkit

GF GeometricFactor

I2C Inter-IntegratedCircuit

JPL JetPropulsionLaboratory

IEM IntegratedElectronicModule

LSB LeastSignificantBit

LSK LeapsecondsKernel(SPICE)

MCP Micro-channelPlate

MESSENGER MErcurySurface,SpaceENvironment,GEochemistry,andRanging

MET MissionElapsedTime

MIDL MissionIndependentDataLayer

MSO Mercury-centricSolarOrbital

NAIF NavigationandAncillaryInformationFacility

NASA NationalAeronauticsandSpaceAdministration

PCK PlanetaryConstantKernel(SPICE)

PDS PlanetaryDataSystem

PHA PulseHeightAnalysis

PPI PlanetaryPlasmaInteractionsPDSNode

RDR ReducedDataRecord

SCLK SpaceClockKernel(SPICE)

SOC ScienceOperationsCenter

 $SPICE \\ Spacecraft, Planet, Instrument, C-matrix Events, refers to the kernel files and NAIF$

software used to generate viewing geometry.

SPK SpacecraftandPlanetsKernel(SPICE)

SSD Solid-StateDetector

SSR SpaceSciencesReview

TOF TimeofFlight

UTC CoordinatedUniversalTime